



# Prompt gamma-ray measurements in ICF experiments

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ACS Symposium, San Francisco, CA

Aug 10, 2014

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# ICF Gamma Ray Physics

a LANL-led collaboration across multiple institutions

## Acknowledgements:

**Y.H. Kim, M. Schmitt,  
N.M. Hoffman, C.S. Young, J.M. Mack,  
D.C. Wilson, S.E. Caldwell, J. Oertel,  
F. Lopez, V. Fatherly, J. Jorgenson,  
A. Hayes-Sterbenz, G. Hale, S.H. Batha**



**C.J. Horsfield, M. Rubery, S. Gales,  
A. Leatherland, W. Garbett**



**T. Hilsabeck, J. Kilkenny**



**J. Milnes, J. Howorth**



**W. Stoeffl, J.A. Church, L. Bernstein,  
D. Sayre, J. Lieberman, A.C. Carpenter,  
D. Casey, C. Cerjan, D. Schneider,  
D. Fortner, A. Mackinnon & the NIF Team**



**A. Zylstra, M. Gatu-Johnson,  
A. J. Frenje, R. Petrasso**



**V. Glebov, W. Shmayda, D. Jacobs-Perkins**



**E.K. Miller, R. Malone, M. Kauffman**



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# Outline

- Intro to Gas Cherenkov Detectors (GCD)
- Gamma Reaction History
- Future Gamma Diagnostics

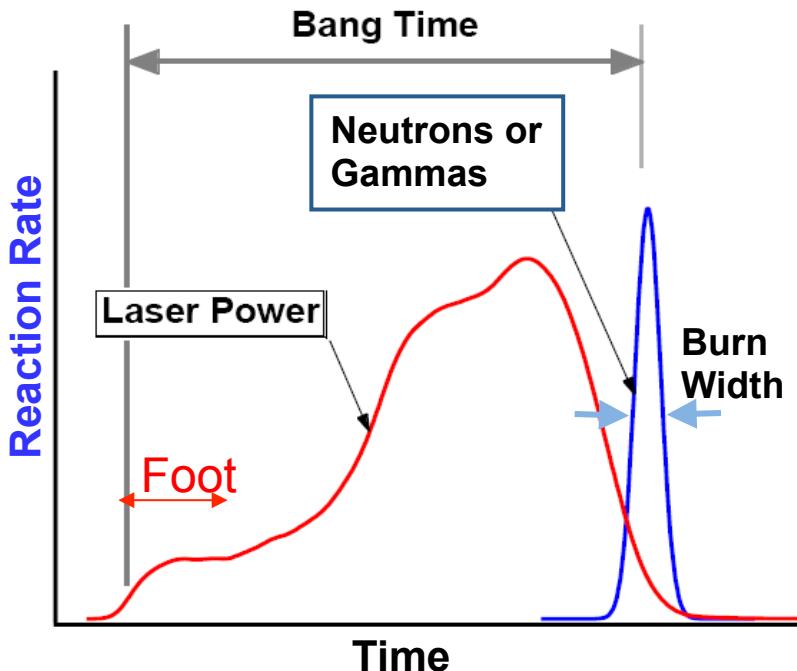
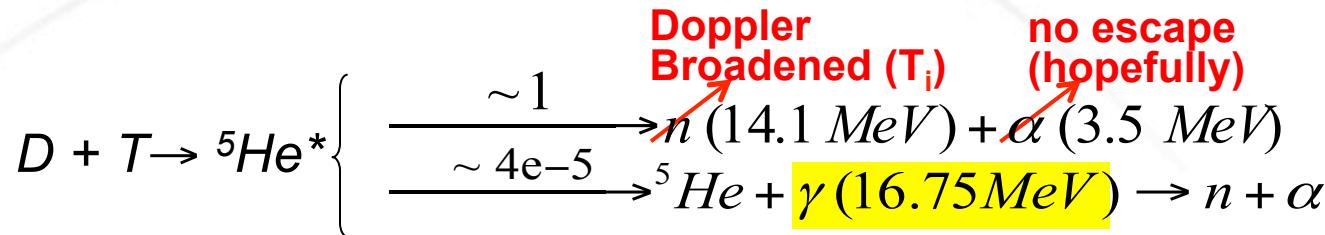
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Operated by Los Alamos National Security, LLC for the U.S. Department of Energy's NNSA



# Gamma-rays provide the most un-perturbed nuclear signature of ICF performance



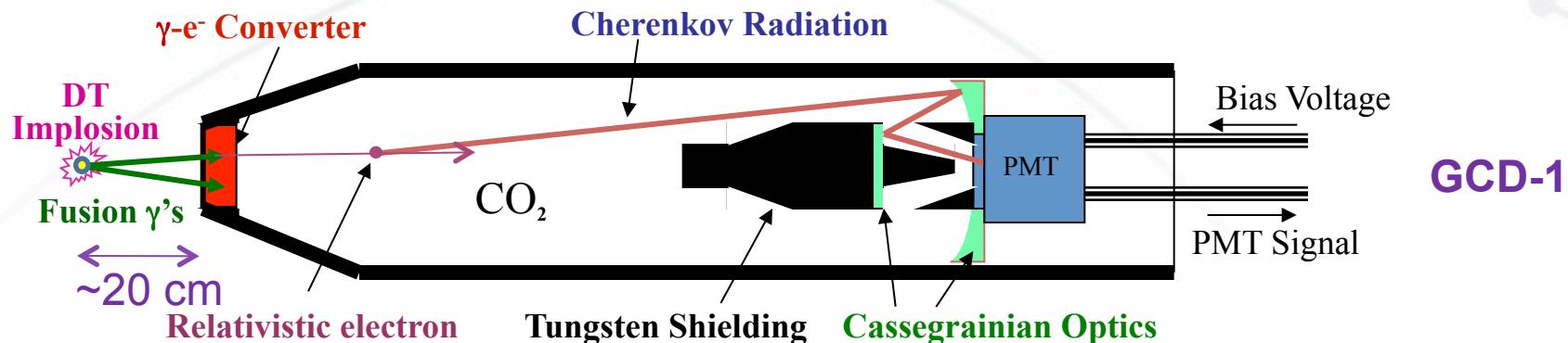
- **Bang Time** - used to establish laser energy coupling to target (shell velocity)

$$ITF = I_0 S^3 \left( \frac{v}{v_0} \right)^8 \left( \frac{\alpha}{\alpha_0} \right)^{-4} \left( 1 - 1.2 \frac{\Delta R_{hotspot}^{K-wtd}}{R_{hotspot}} \right)^{4+\epsilon} \left( \frac{M_{clean}}{M_{DT}} \right)^{0.5}$$

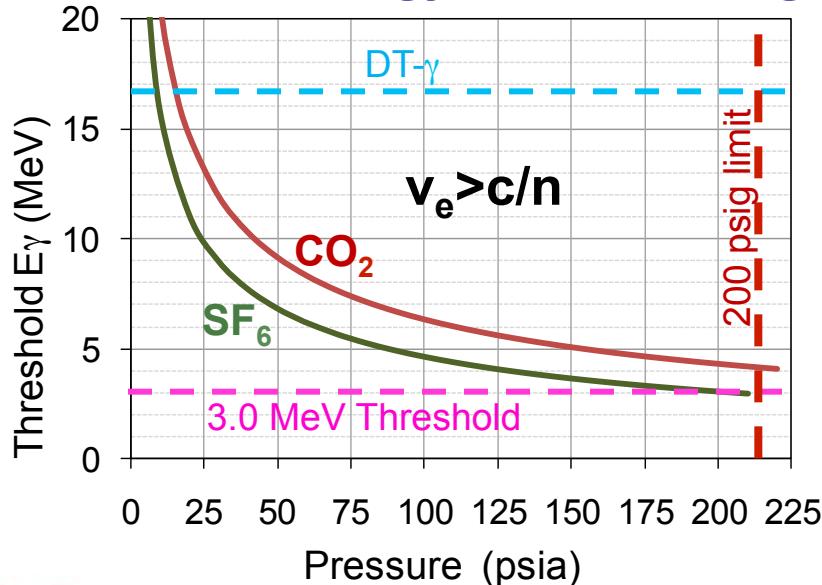
- **Burn Width & other RH features** - used for failure mode correlation

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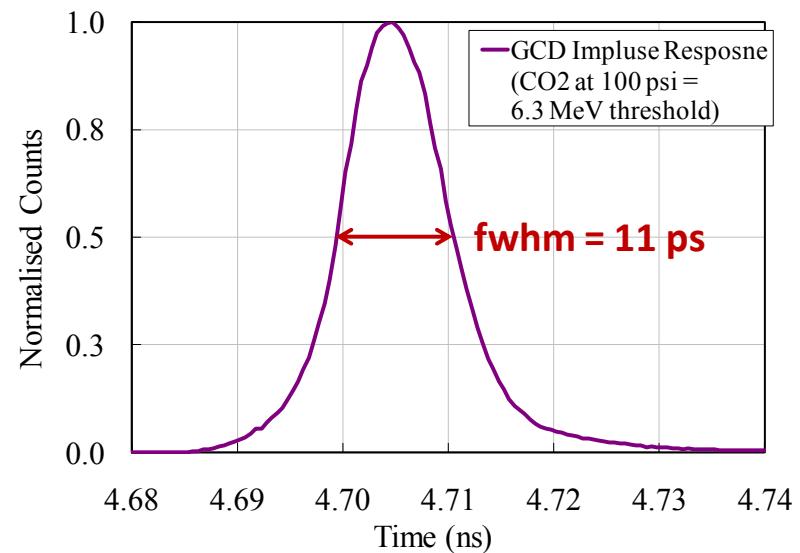
# Gas Cherenkov Detectors (GCD) convert MeV gammas to UV/Visible for easy detection



## Variable Energy Thresholding

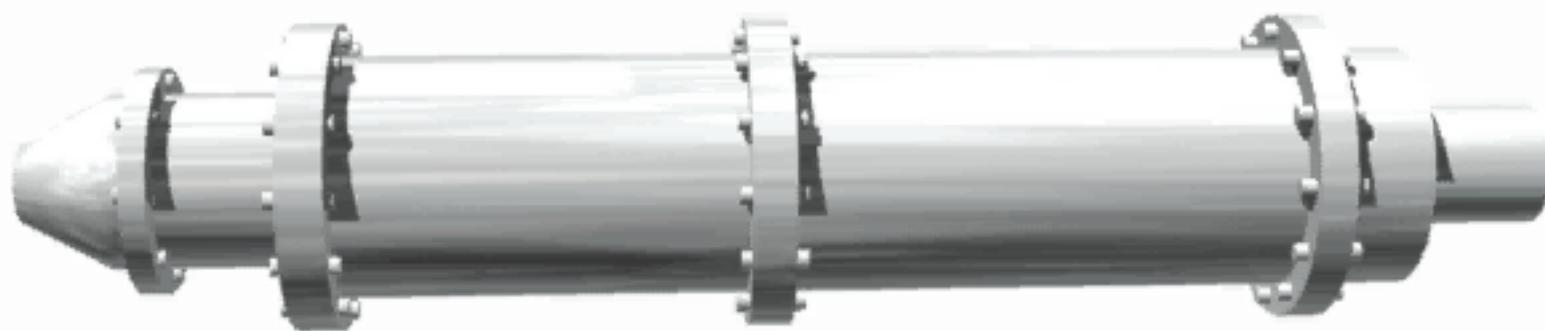


## Fast Time Response



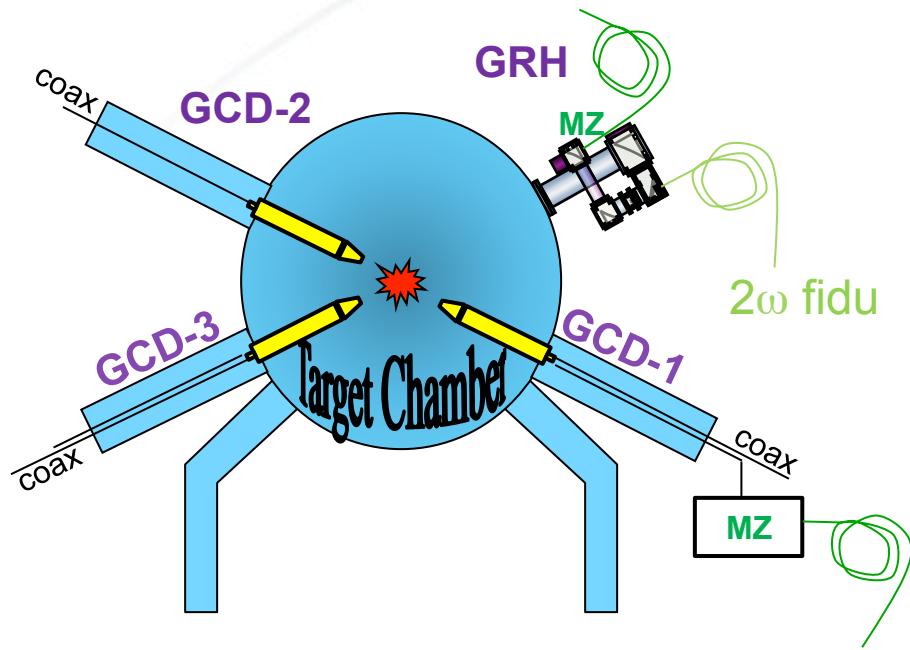
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# GCD Animation

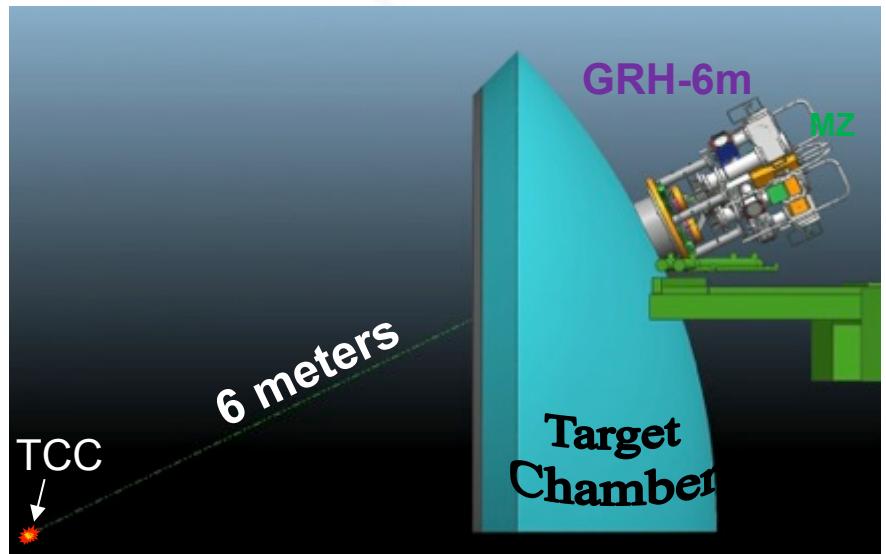


# Gamma Ray Diagnostic capability at OMEGA & NIF

OMEGA-60



NIF

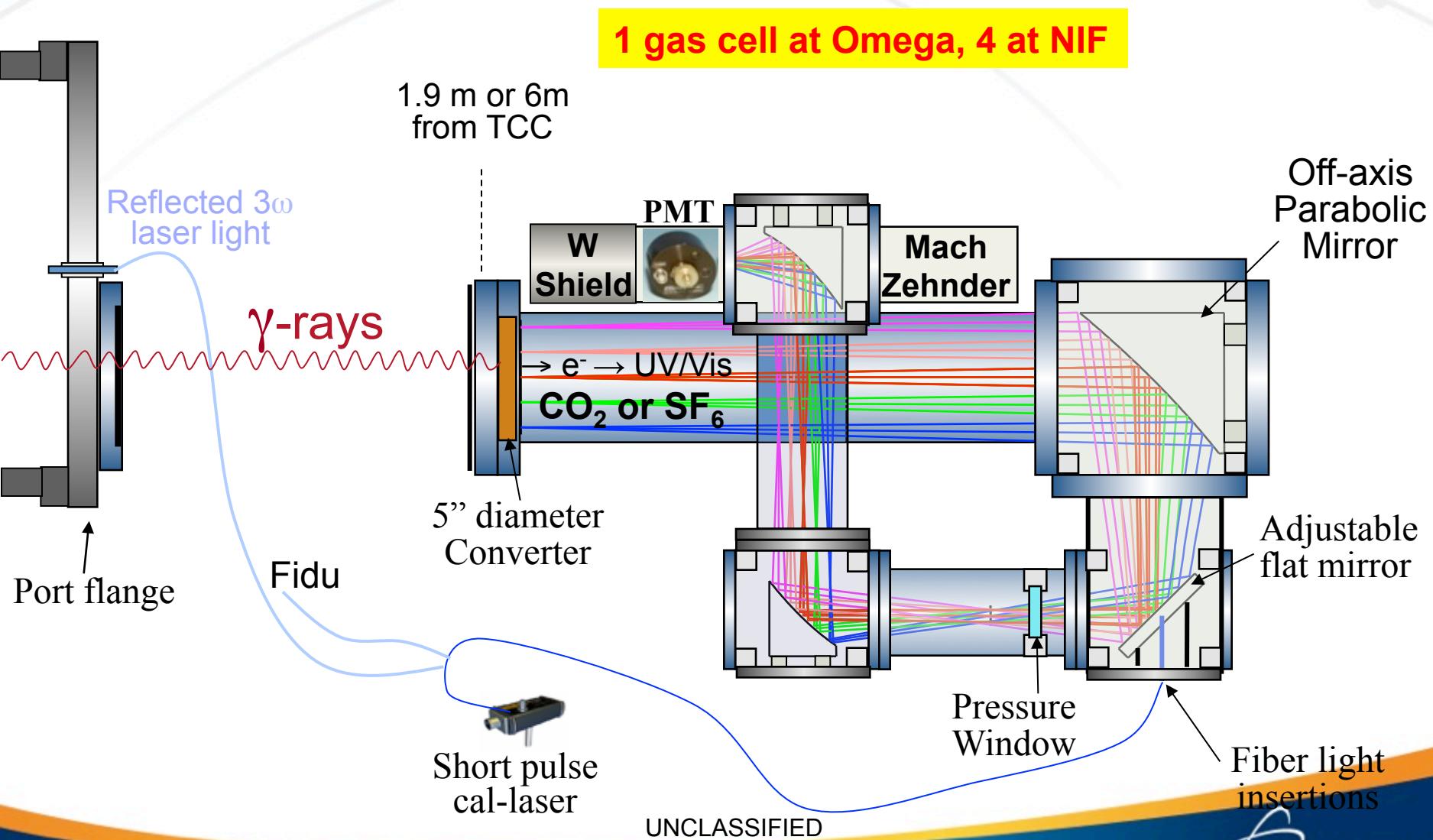


- 3 GCDs (20cm), 1 GRH (187cm)
  - Coax (40ft) & Mach Zehnders
  - Only GRH absolutely timed

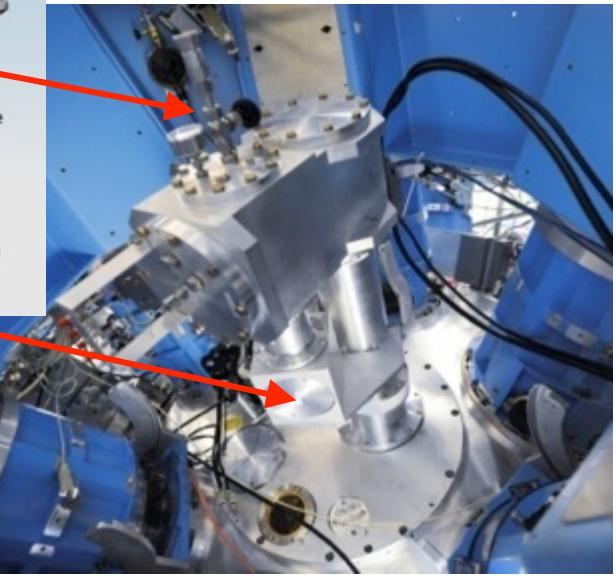
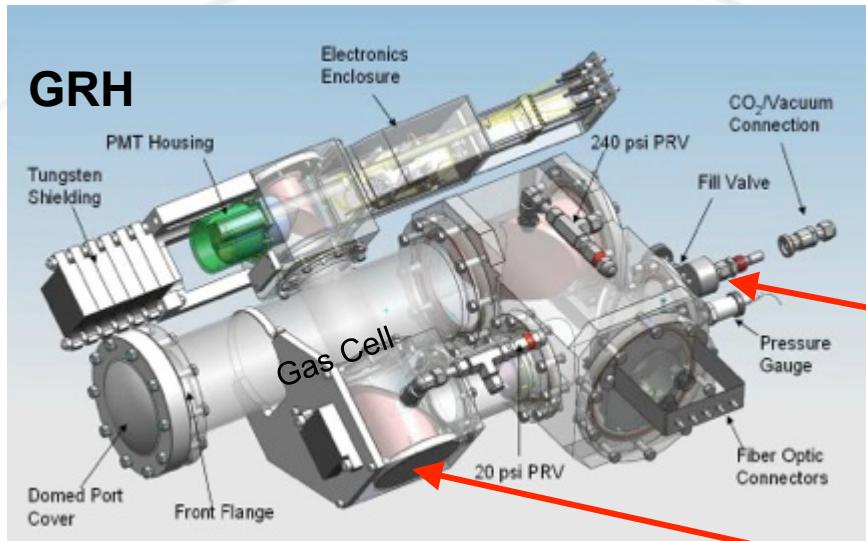
- 4 GRHs (607 cm)
  - Mach Zehnders only (160 ft)
  - All absolutely timed

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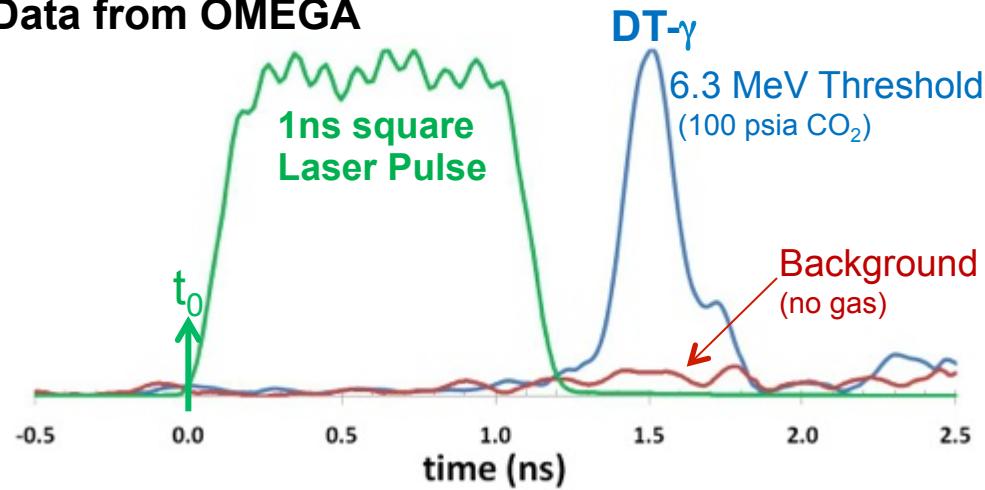
# Gamma Reaction History (GRH) is optimized to operate outside target chamber.



# Single-channel GRH prototype performance demonstrated at OMEGA (U. of Rochester) in 2009



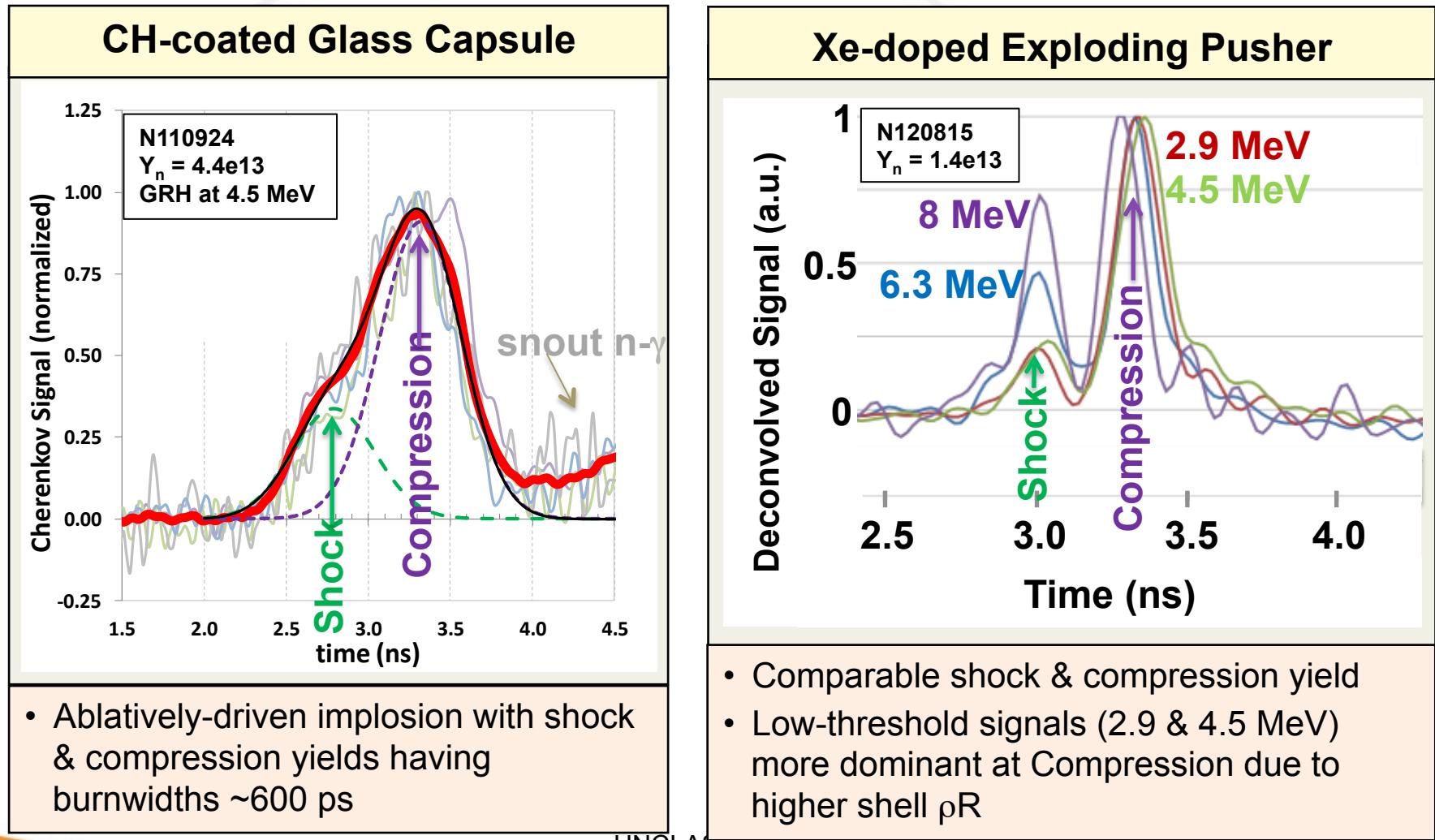
Data from OMEGA



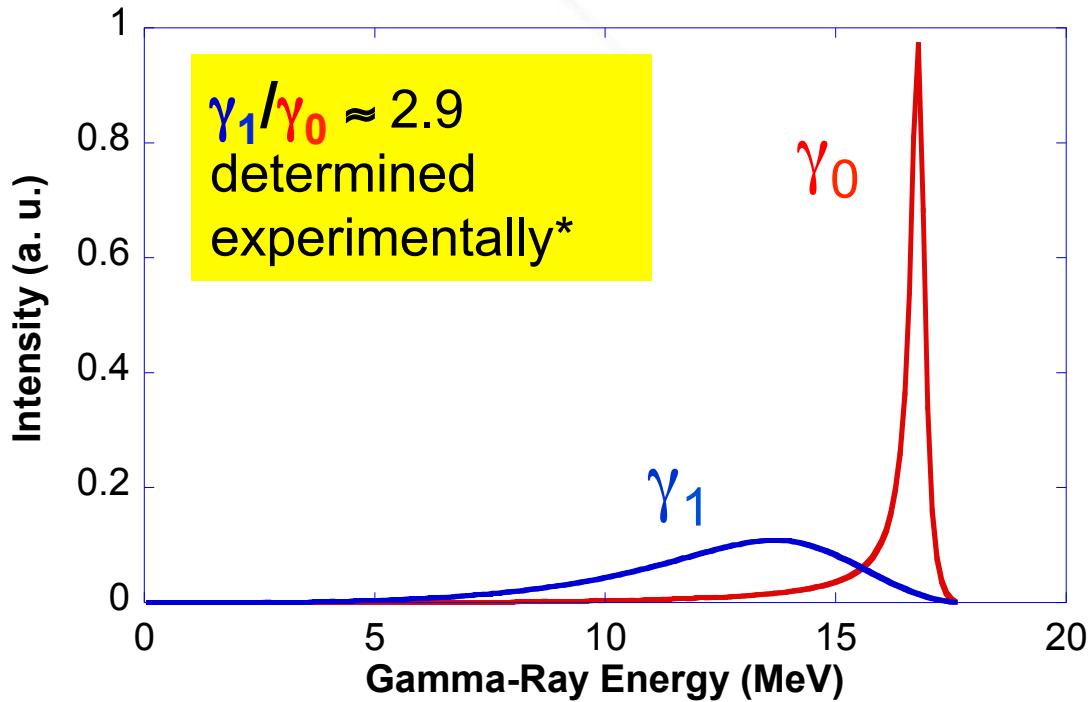
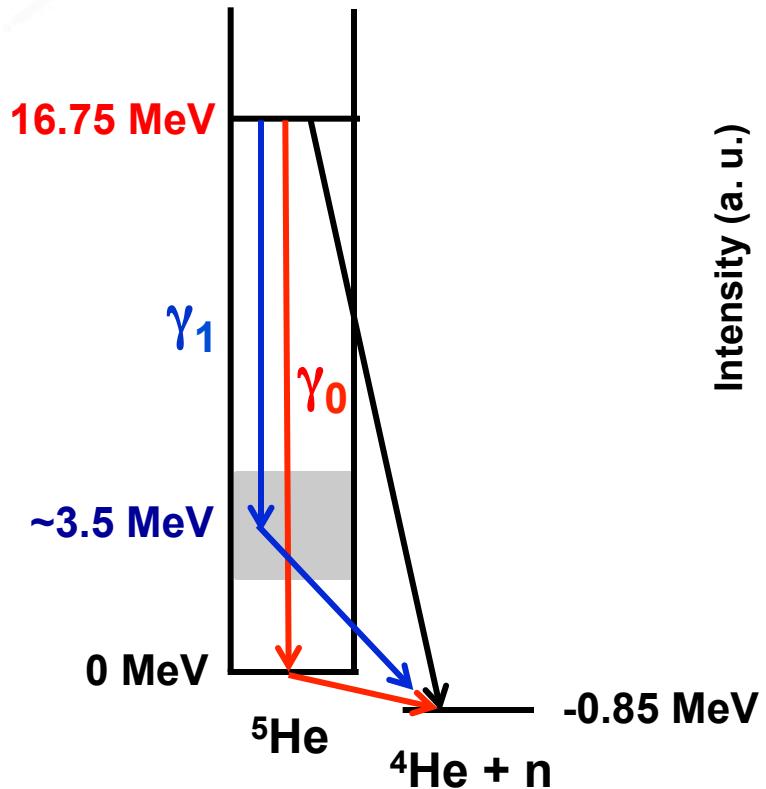
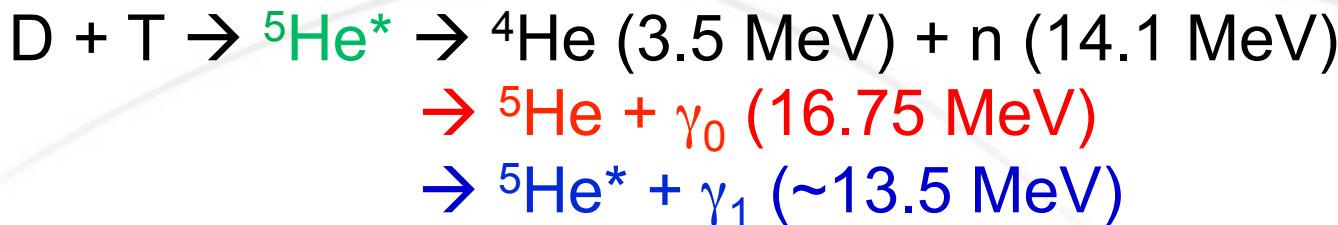
Installed on OMEGA

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# NIF has produced Interesting Direct-Drive Reaction Histories



# DT $\gamma$ -ray Spectrum consists of 2 prominent lines



\* Courtesy C.J. Horsfield (AWE),  
assuming lines shapes offered by  
R-Matrix analysis (G. Hale, LANL)

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# GCD & GRH have provided great HED & Burn Physics results at OMEGA

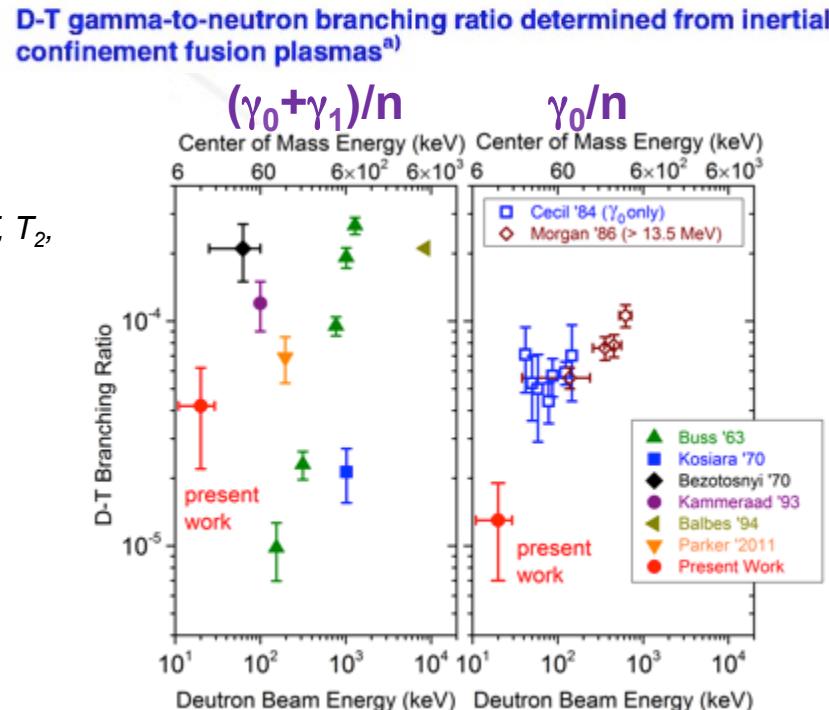
PHYSICS OF PLASMAS 19, 056313 (2012)

## Accomplished:

- DT Branching Ratio =  $(4.2 \pm 2)e-5 \gamma/n$
- Characterization of other fusion gammas ( $D^3He$ , HT,  $T_2$ ,  $T_3He$ ,  $^3He^3He$ , HD,...)
- $(n,n')$  gammas from pucks of various materials → ablator areal density of CH & SiO<sub>2</sub> implosions

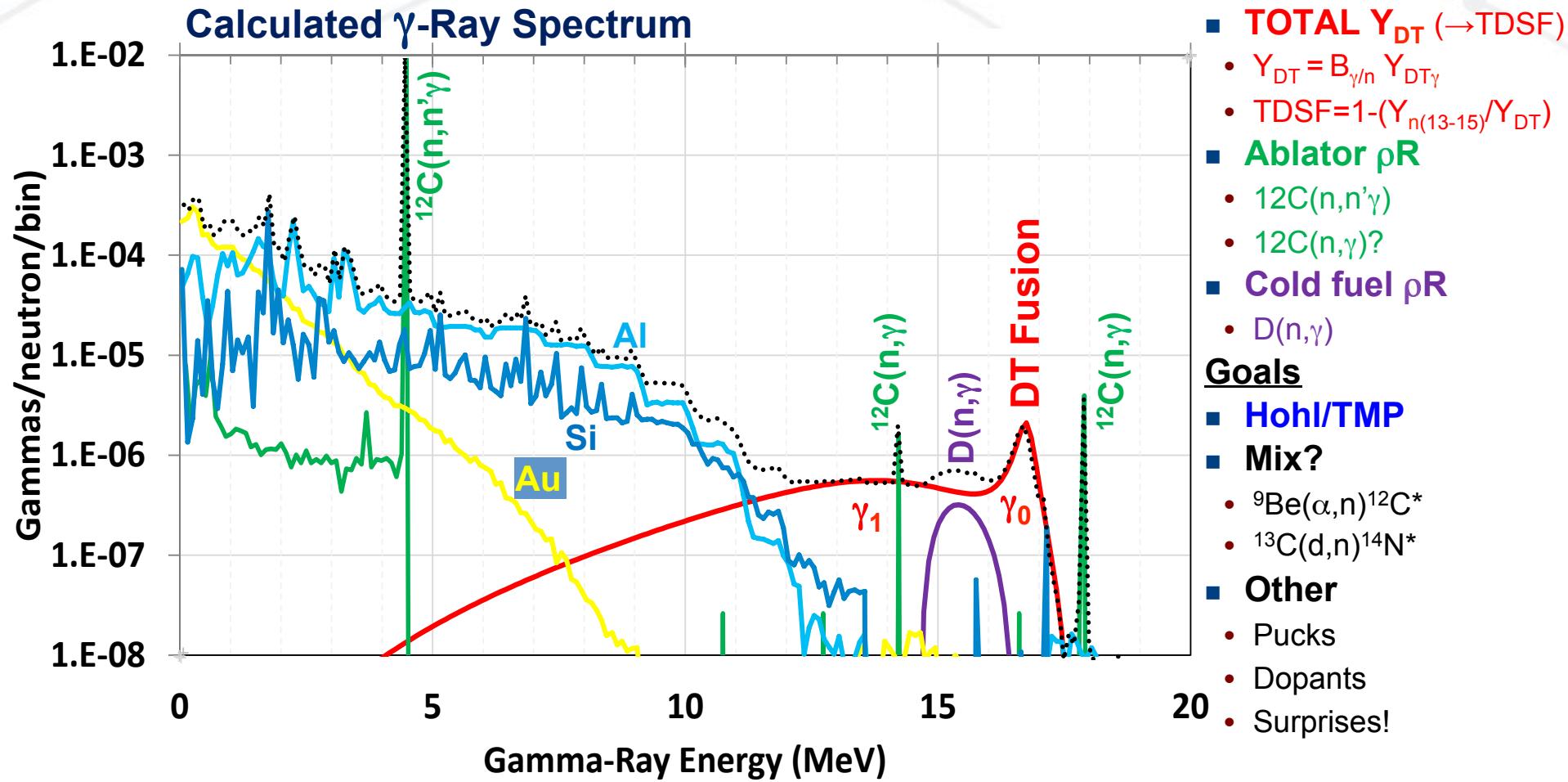
## In Progress:

- Kinetic Plasma Effects
  - Fuel Ion Segregation
  - Knudsen Reactivity Reduction
  - Transport validation (mass, momentum, energy)
- Charged-Particle Stopping Power
- Charged-Particle induced gammas for Mix diagnosis

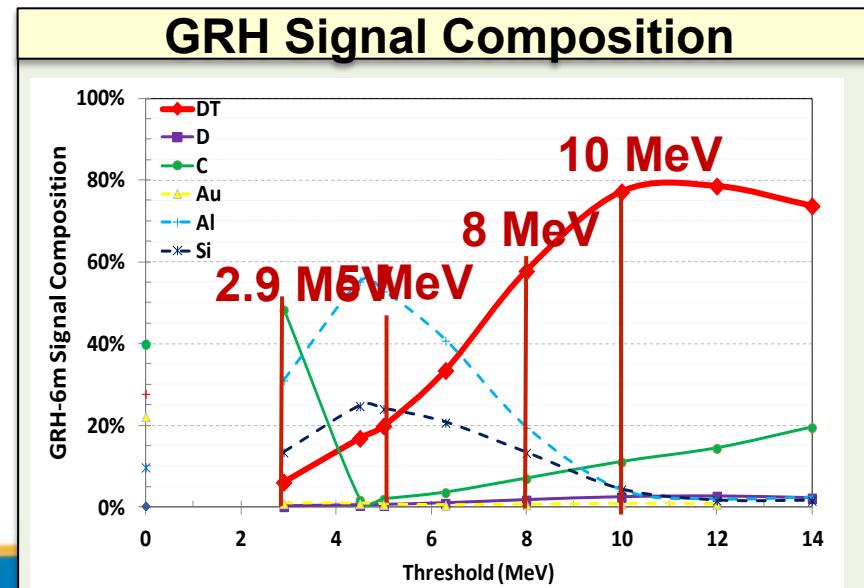
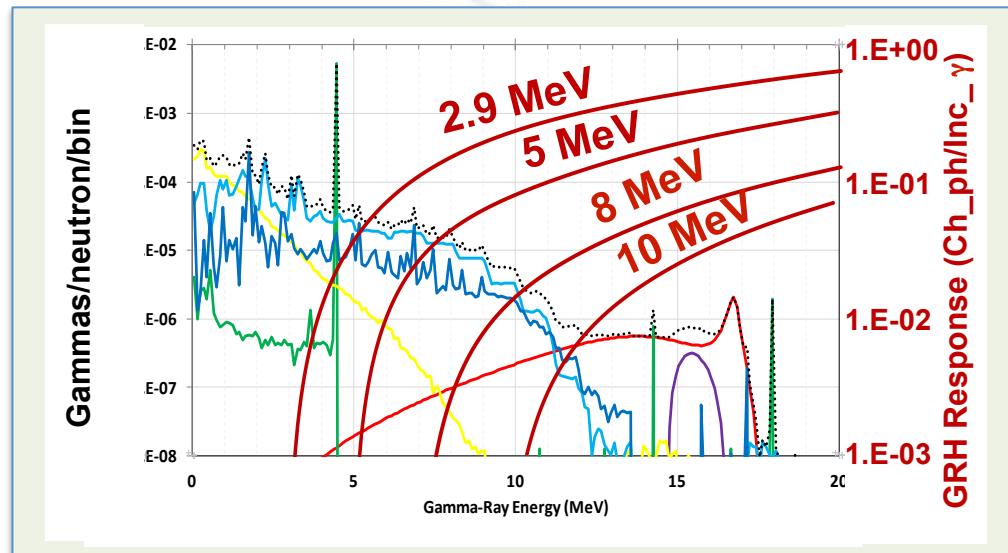
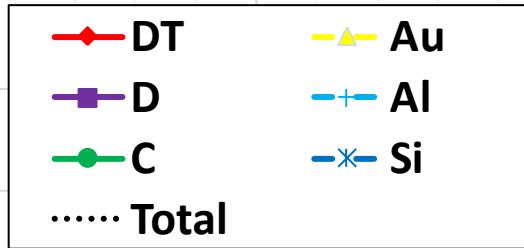


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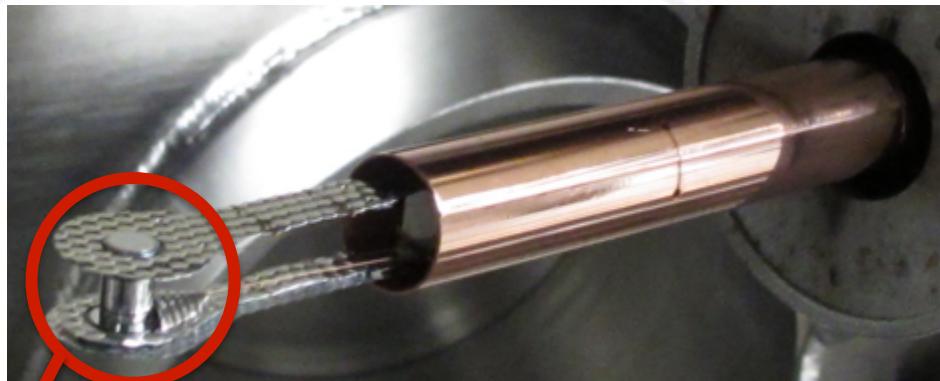
# The Prompt $\gamma$ -Ray Energy Spectrum from Indirect-Drive, Cryo-Layered Implosions is full of information!



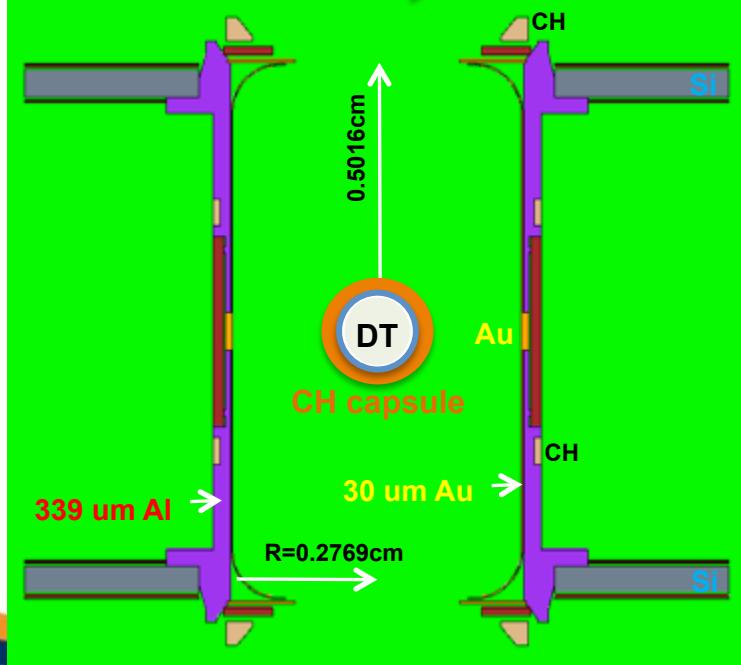
# GRH isolates DT fusion and $^{12}\text{C}(\text{n},\text{n}')$ $\gamma$ -rays



# 14 MeV neutron-induced $\gamma$ -Rays from CH Capsule & Hohlraum assembly are simulated in MCNP

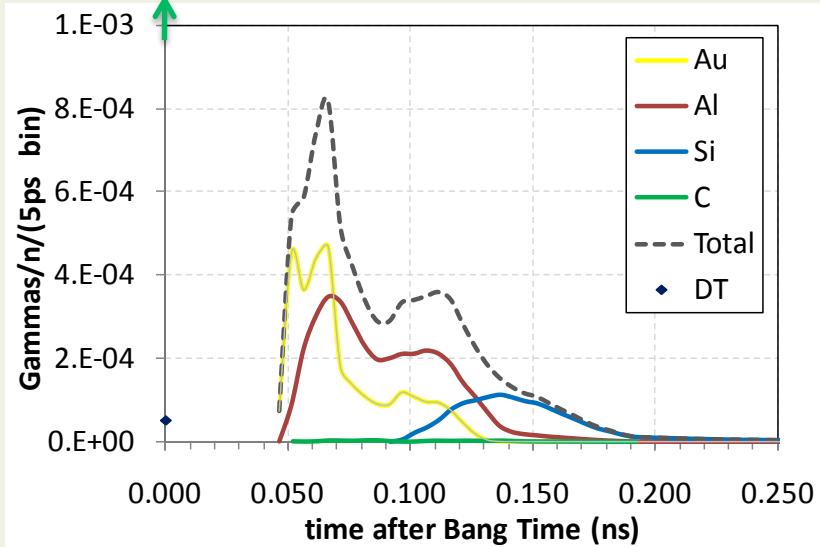


544 Hohlraum

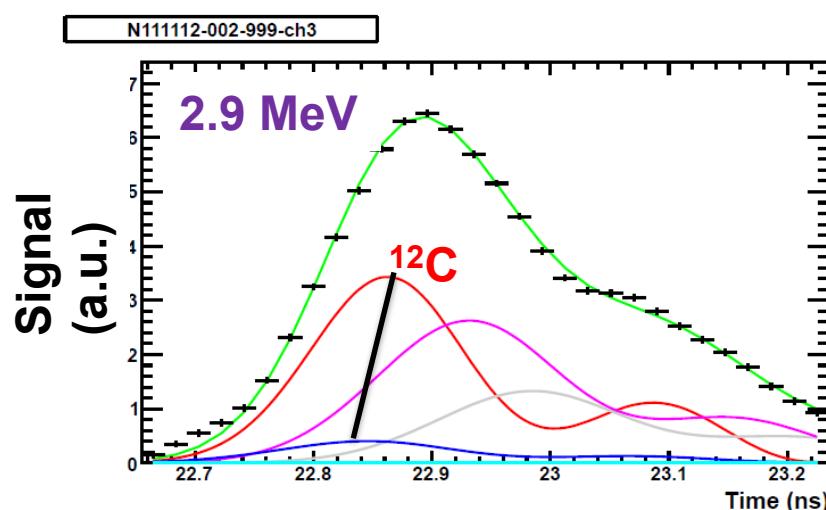
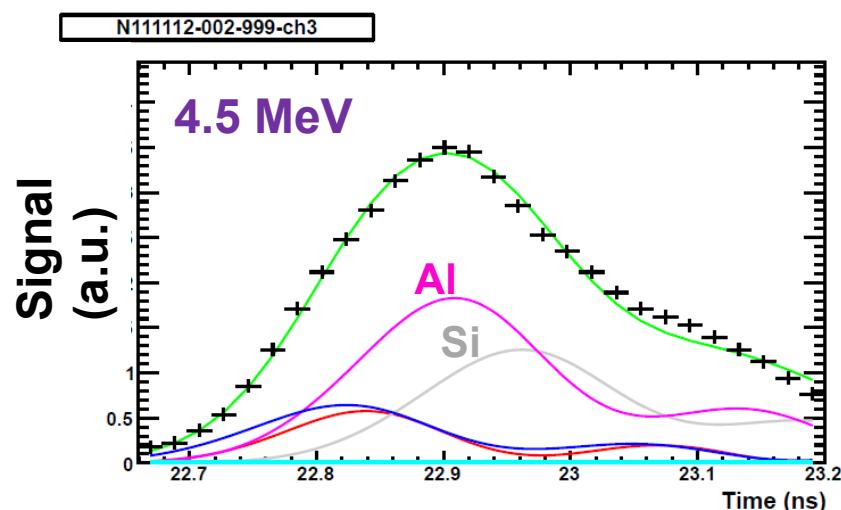
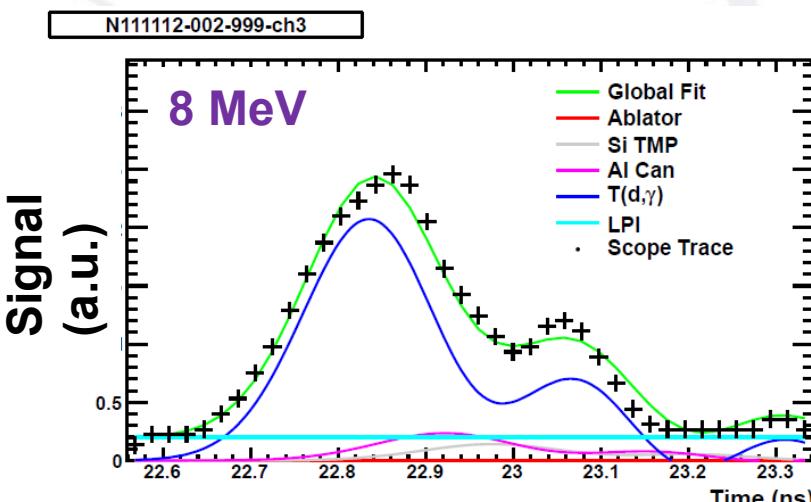
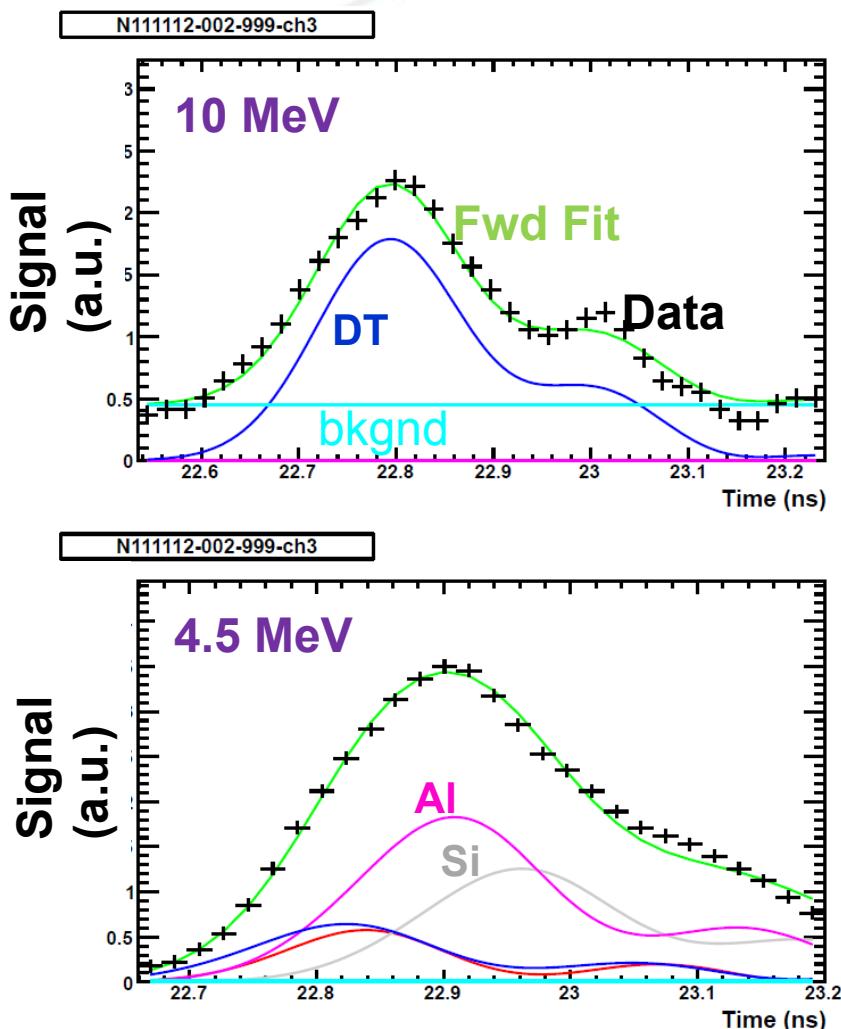


Prompt  $\gamma$ -Ray Temporal History

BT+ $\frac{1}{4}$  ns,  $\delta(t=0)$  14 MeV n-source



# Gaussian forward fit decomposition into spectral components provides Total DT & $^{12}\text{C}(\text{n},\text{n}'\gamma)$ yields and absolute timing (BT, BW, $t_{\text{C}\gamma}$ )

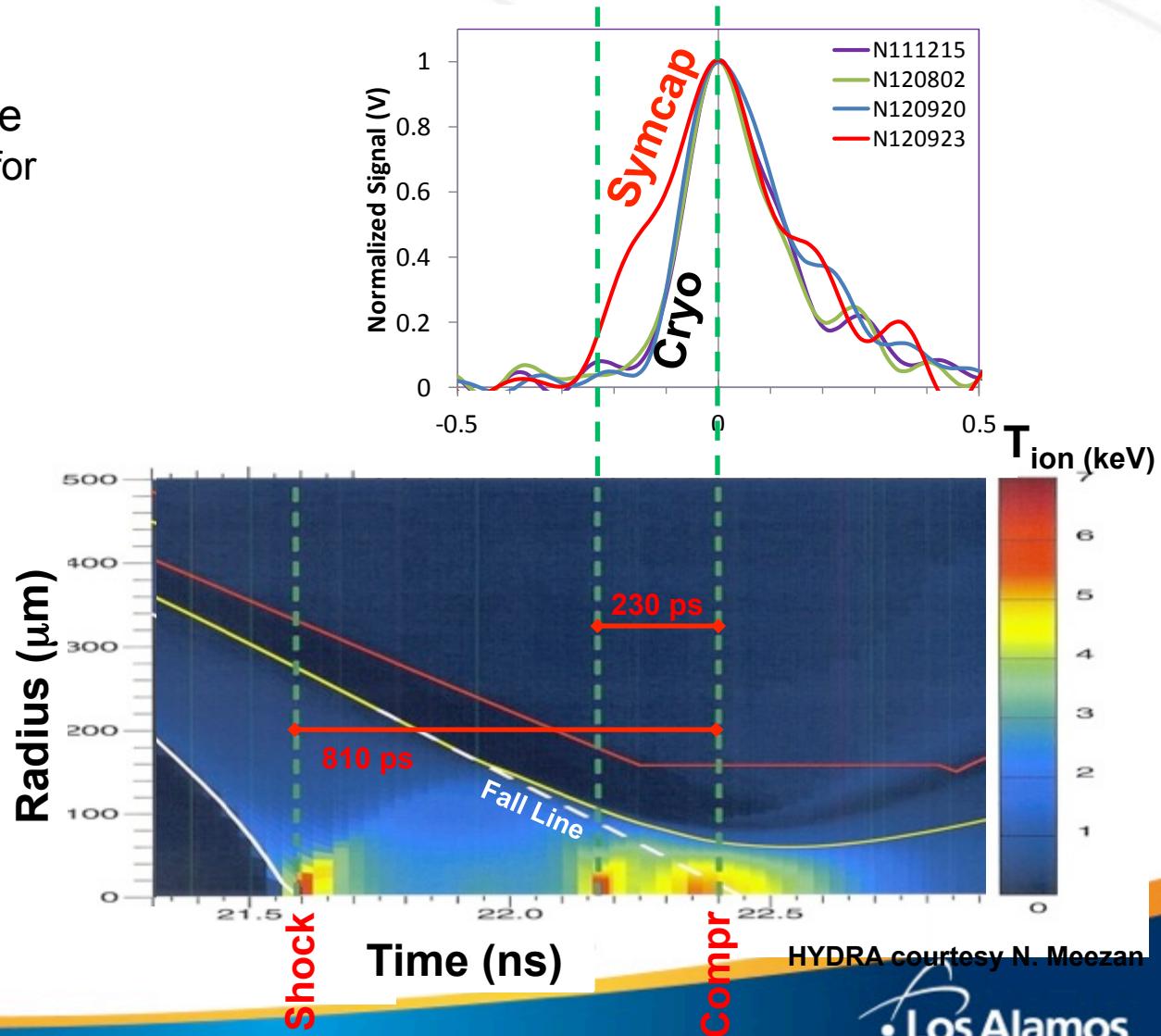
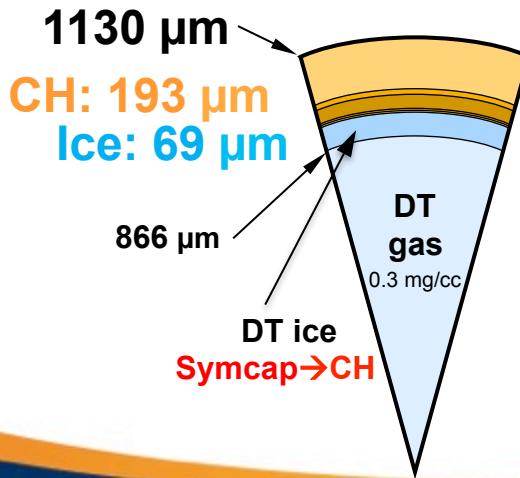


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$t_{\text{C}\gamma} > \text{BT}$

# Indirect Drive: DT Symcap produces discernible “Re-Shock flash”, Cryo-layered implosions do not

- Symcap uses surrogate CH layer in place of DT ice
  - provides harder boundary for shock reflection
- Re-Shock flash occurs before Fall Line reaches core
  - $Y_{\text{Compr}} / Y_{\text{Re-Shock}}$  is a measure of mix

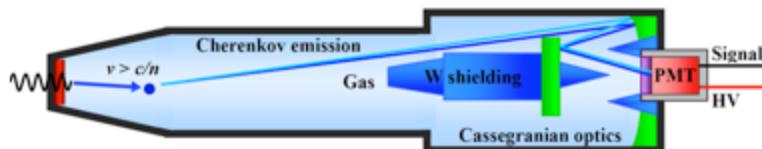


HYDRA courtesy N. Meezan  
• Los Alamos  
NATIONAL LABORATORY  
EST. 1943

# Future gamma diagnostics will add significant capability

- Gas Cherenkov Detectors (temporal detectors):

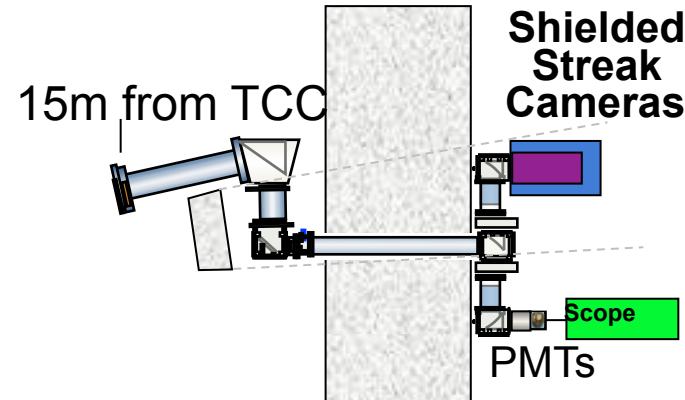
## “Super” GCD



400 psia  $\text{C}_2\text{F}_6$

- Low Threshold, High Sensitivity
  - ~2 MeV threshold
  - 20 cm from TCC (TIM mounted)

## GRH-15m



- High Temporal Resolution
  - Streak Camera for ~10ps resolution
  - Camera behind 6' bio shield wall

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# “Super” GCD (GCD-3 at $\Omega$ ) provides High-Sensitivity, Low-Threshold capability now at OMEGA and eventually at NIF

## “Super” GCD



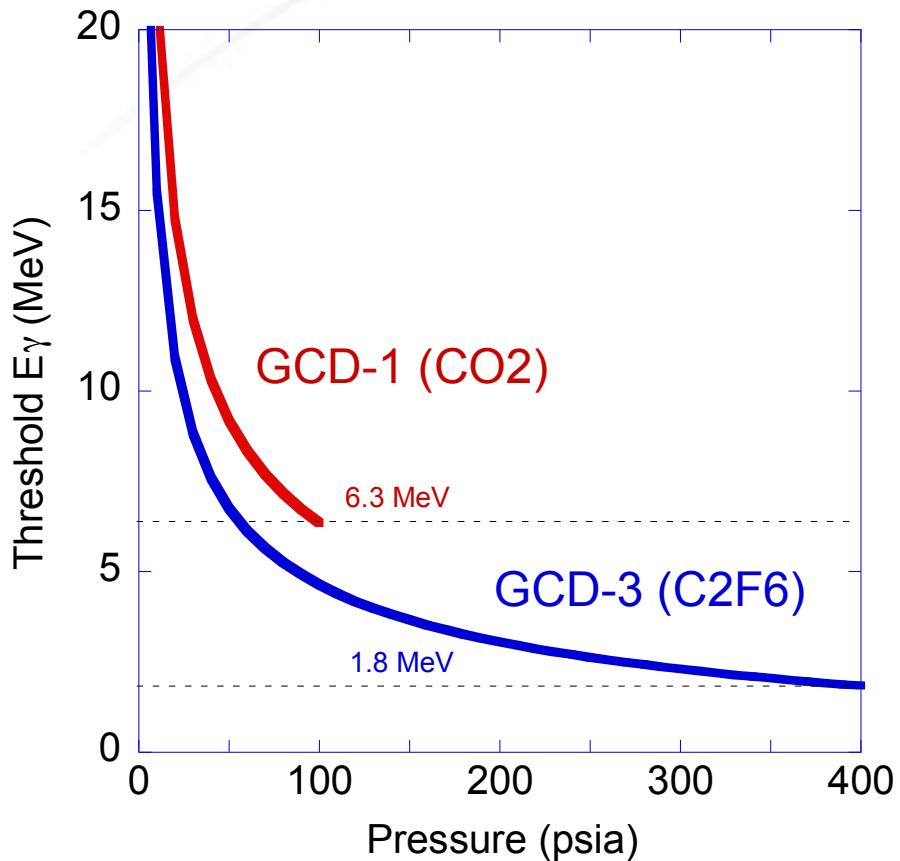
- Low Threshold, High Sensitivity
  - ~2 MeV threshold
  - 20 cm from TCC (TIM mounted)

### Physics Driven Requirements:

- Low Threshold ( $\leq 2$  MeV) to reveal new portions of gamma-ray spectrum
  - High pressure (400 psia) → redesigned pressure boundary
  - Fluorinated gases → metal seals to achieve  $< 1e-9$  scc/s leak rate to avoid damage to TRS catalyst
- High Sensitivity
  - TIM-based to capture solid angle
  - Modular optics package to optimize SNR
- Absolute Timing & Dry Run capability
  - $2\omega$  fidu injection
- Improved SNR
  - better shielding
  - additional precursor to signal delay

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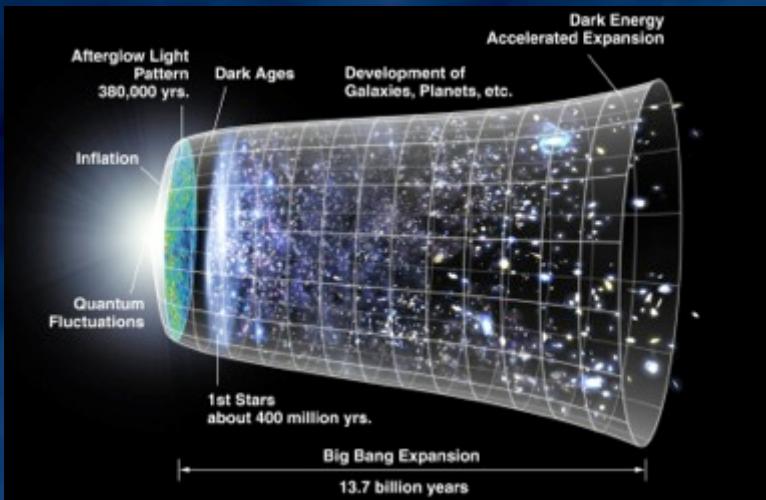
# Lower Energy Threshold (~ 2 MeV) opens up new portions of gamma-ray study



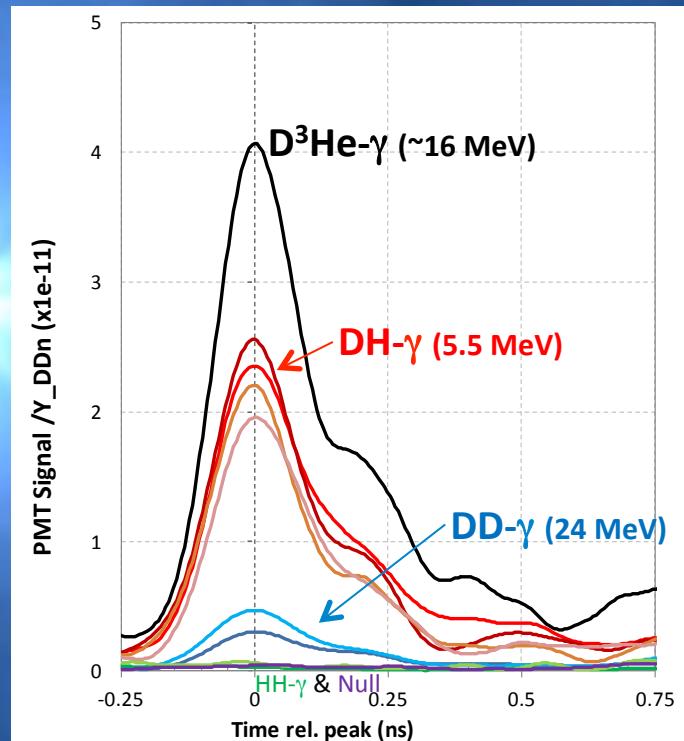
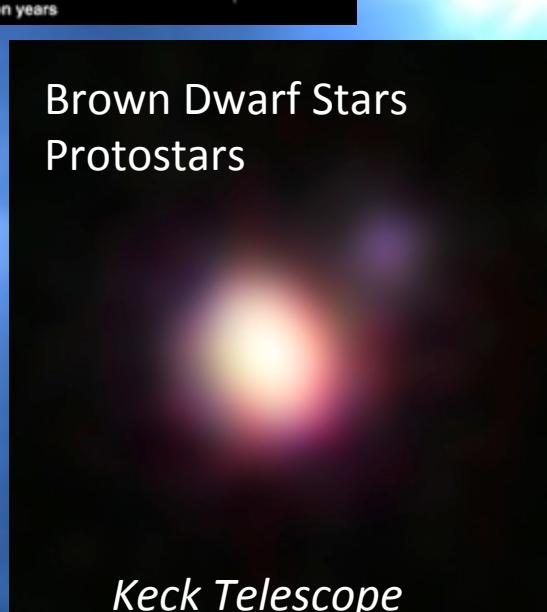
- New gamma-ray detection (too low E for GCD-1, too dim for GRH):
  - $^{16}\text{O}(\text{n},\text{n}'\gamma)$  at **6.1 MeV** ( $\text{SiO}_2 \rho\text{R}$ )
  - $^{13}\text{C}(\text{d}_{\text{k}\alpha},\text{n})^{14}\text{N}^*$  at **5.69 MeV** (CH Mix)
  - $^{9}\text{Be}(\alpha,\text{n})^{12}\text{C}^*$  at **4.44 MeV** (Be Mix)
  - $^{9}\text{Be}(\text{d}_{\text{k}\alpha},\text{n})^{10}\text{B}^*$  **3.4 MeV** (Be Mix)
  - $^{10}\text{B}(\text{d}_{\text{k}\alpha},\text{n})^{11}\text{C}^*$  at **~7 MeV** ( $\text{B}_4\text{C}$  or BH Mix)
  - HD- $\gamma$  at **5.5 MeV** (MIT Zylstra PhD)

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# The p+D reaction relevant to BBN, formation of proto stars, and brown dwarfs, is being investigated at OMEGA in July-August

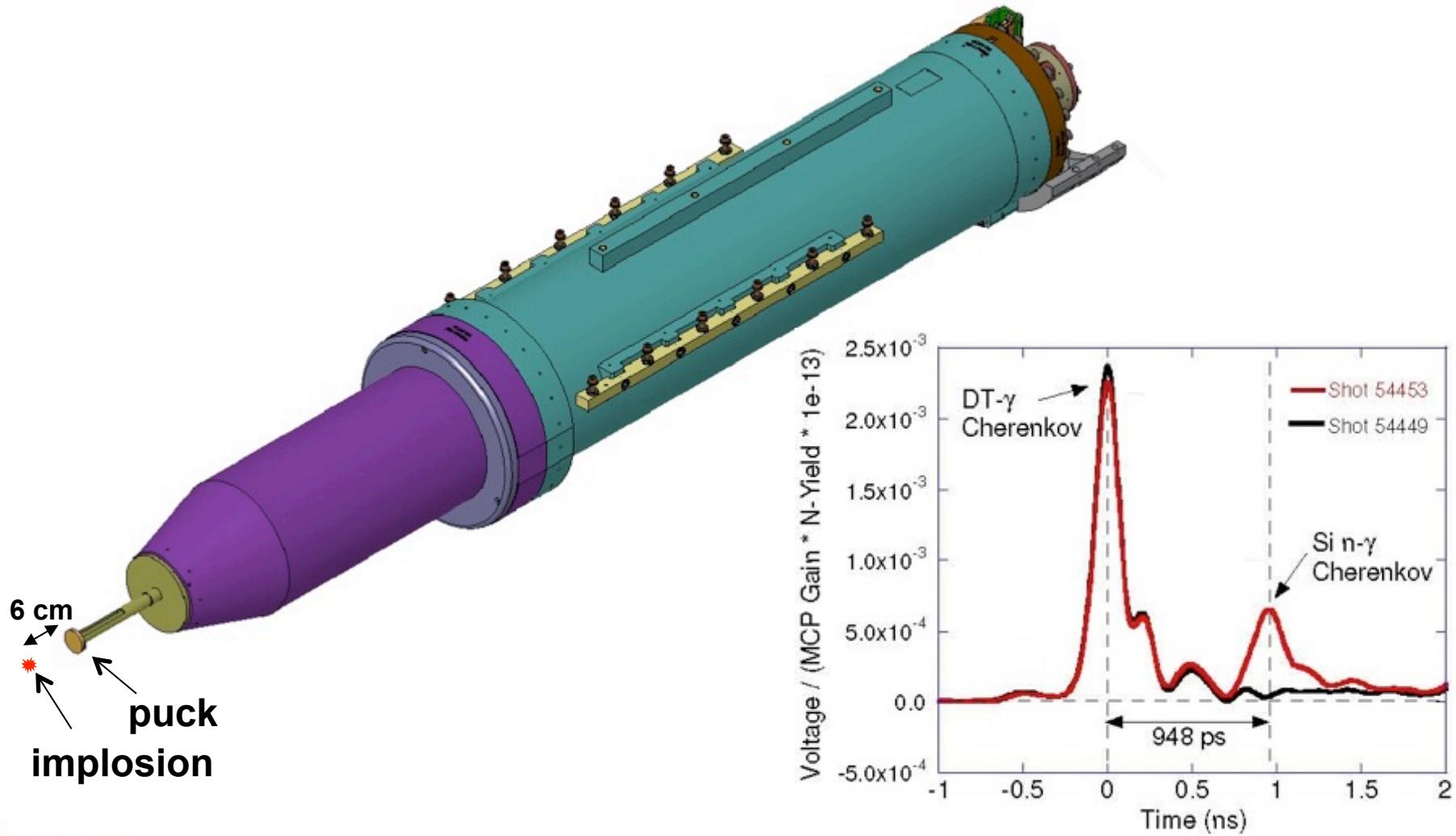


A. Zylstra  
(PhD thesis)



LANL's new GCD-3 uniquely identified HD fusion gammas at 5.5 MeV for the first time (as well as  $D^3\text{He}-\gamma$  &  $D_2-\gamma$ )

# Integral puck holder allows study of 14 MeV neutron interactions with materials placed near implosion

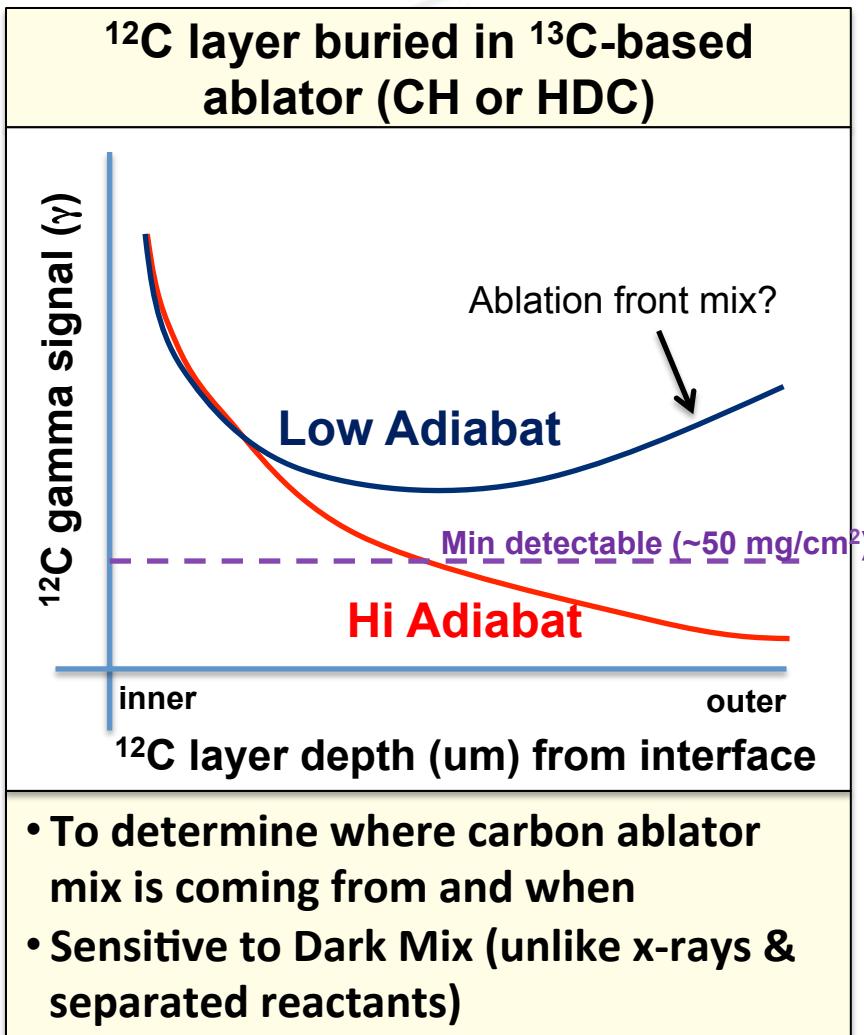


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Y. Kim, et. al, JoP, 2010

# Gamma Rays may illuminate “Dark Mix”

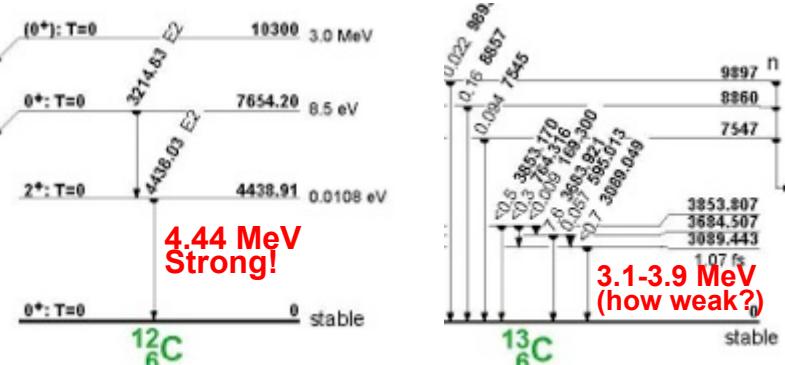
$^{12}\text{C}$  vs  $^{13}\text{C}$  pucks at OMEGA will determine feasibility



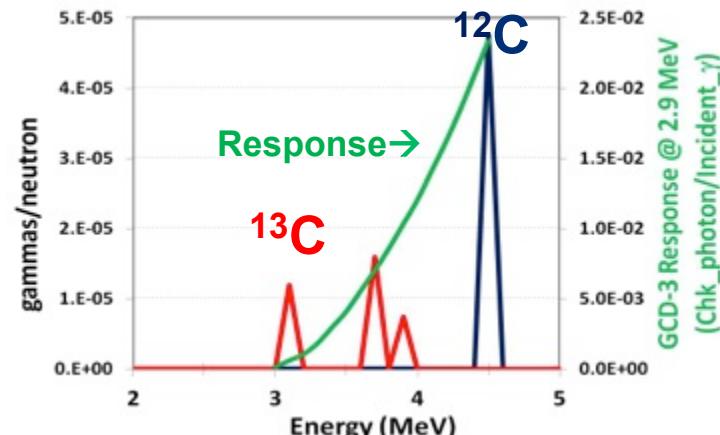
Concept by W. Stoeffl

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## Carbon Energy Level Diagrams

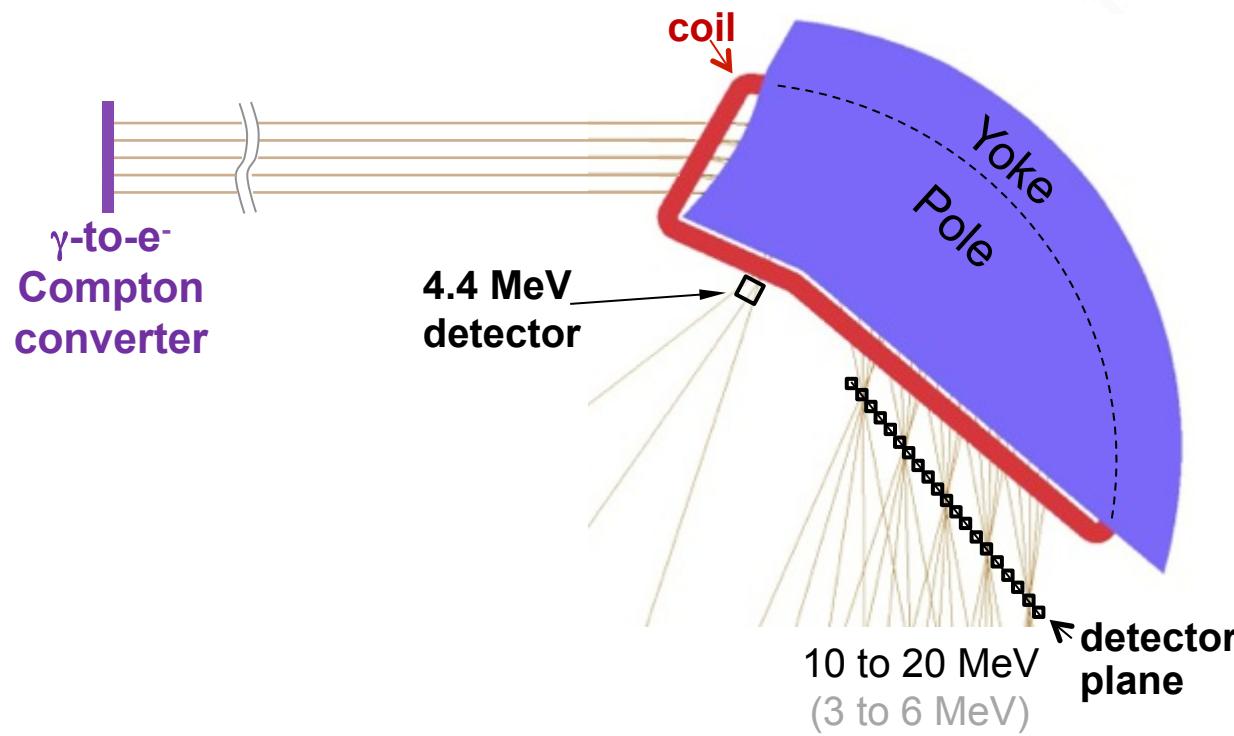


MCNP predicts ~6x less signal for  $^{13}\text{C}$ , but large uncertainty



# Future gamma diagnostics will add significant capability

- Spectroscopic detectors:
  - Gamma-to-Electron Magnetic Spectrometer (GEMS)



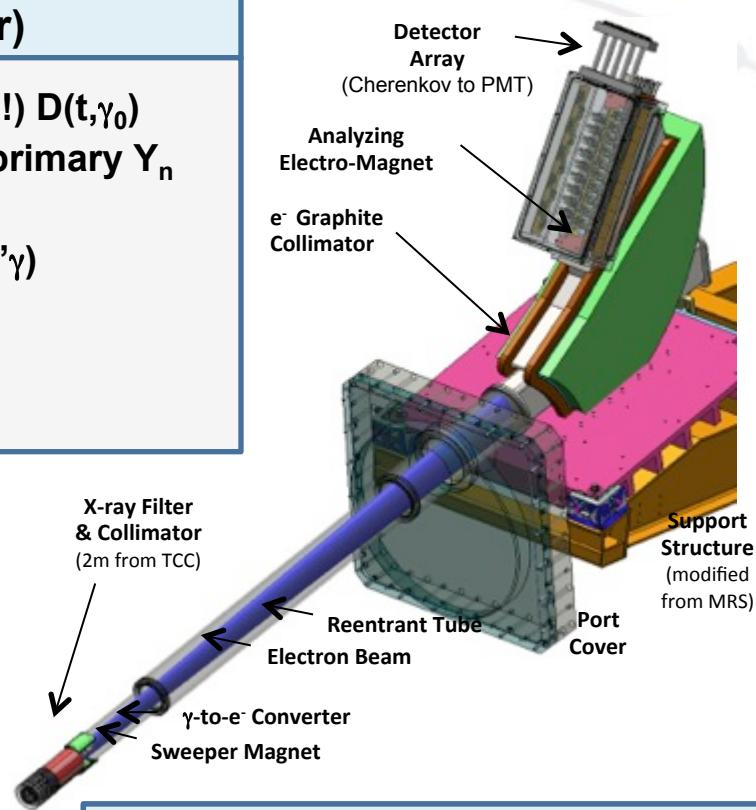
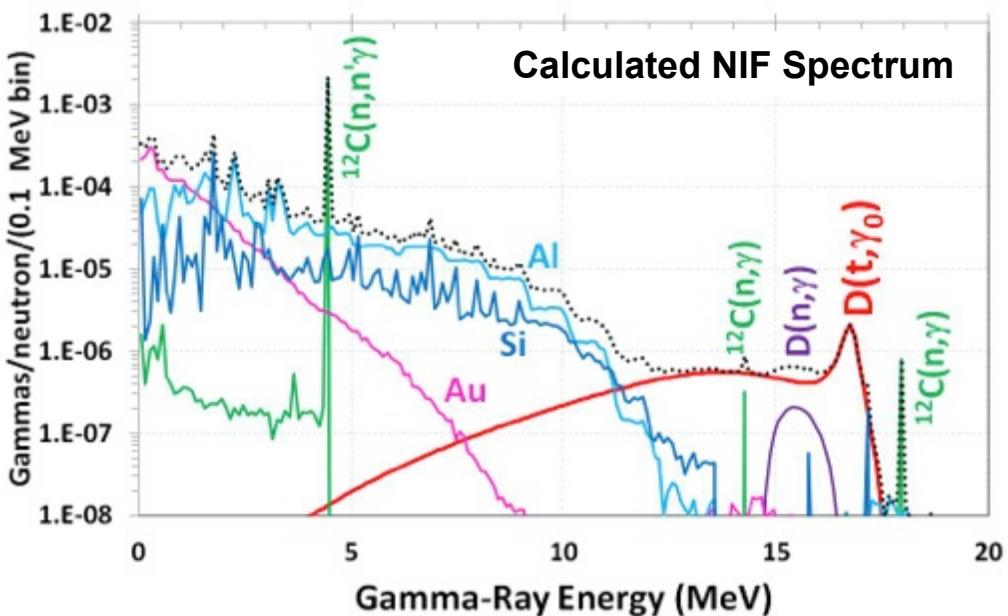
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# Gamma Spectroscopy will be an enabling technology for NIF

## GEMS (Gamma-to-Electron Magnetic Spectrometer)

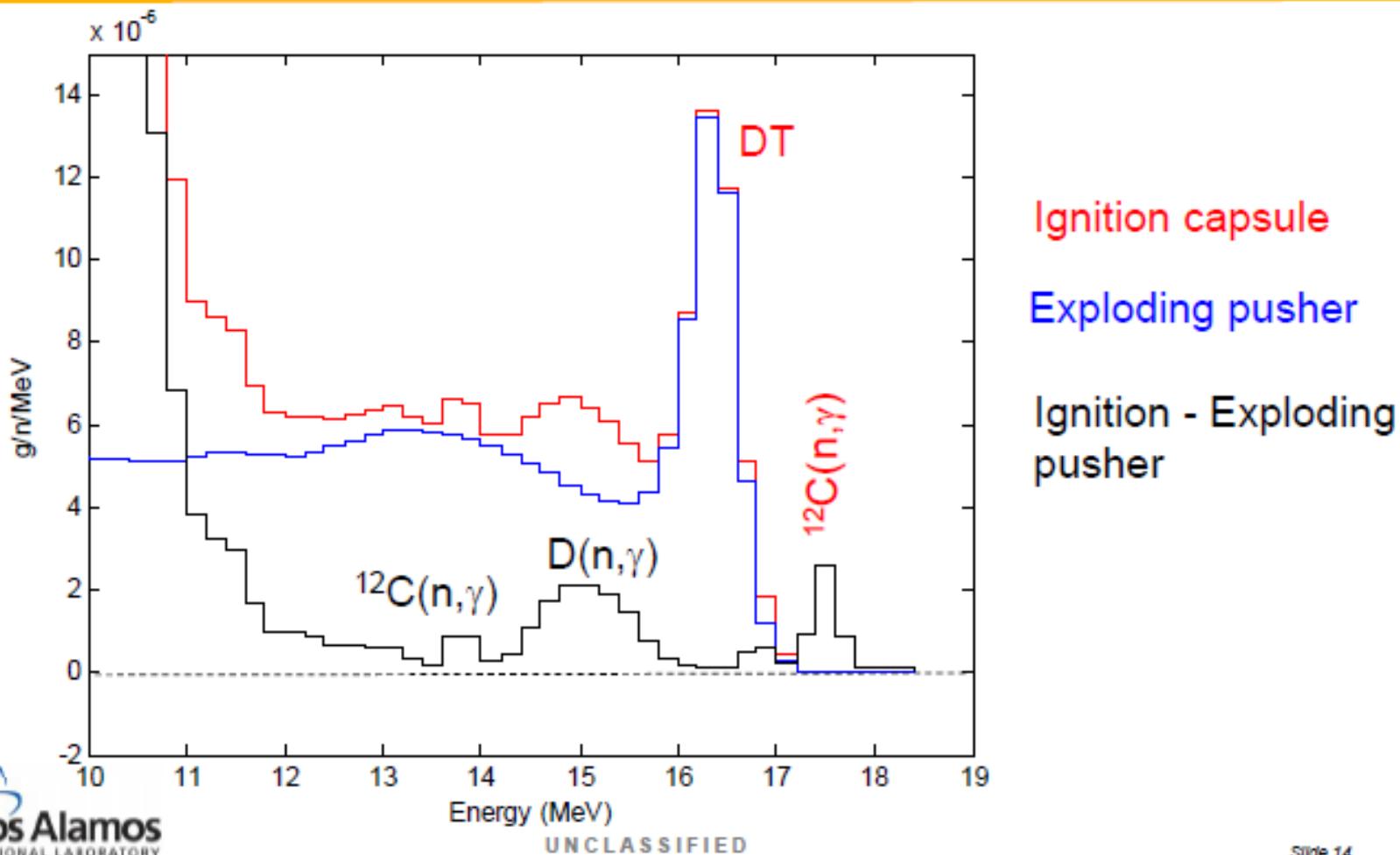
- Total Yield (no Total Yield measurement currently exists!)  $D(t, \gamma_0)$
- Total Down Scatter Fraction (TDSF) when combined w/ primary  $Y_n$
- $4\pi$  Global Fuel  $\rho R$  (Fuel  $\rho R$  currently line-of-sight)  $D(n, \gamma)$
- Ablator  $\rho R$  (reduced uncertainty relative to GRH)  $^{12}\text{C}(n, n'\gamma)$
- Mix studies (e.g.,  $^9\text{Be}(\alpha, n\gamma)$ ,  $^{13}\text{C}(d_{ko}, n\gamma)$ ,  $^{11}\text{B}(d_{ko}, n\gamma)$ )
- Neutron Interactions on materials (i.e., pucks)
- Astrophysical studies (e.g., s & r-processes)



## NIF CDR held in July 2013

- Energy Resolution:  $\Delta E/E \leq 5\%$
- Energy Range:  $E_0 \pm 33\%$  within 2-25 MeV
- Temporal response  $< 1.5$  ns
- Viable at  $Y_{DTn} \geq 5 \times 10^{14}$

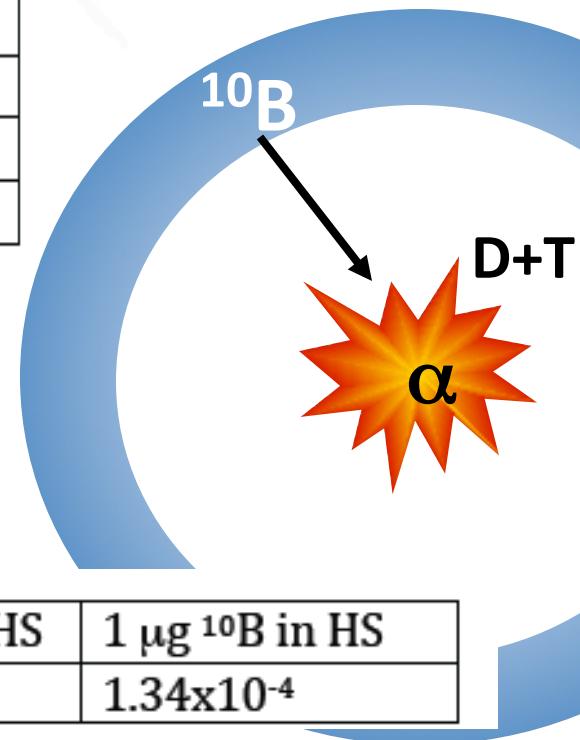
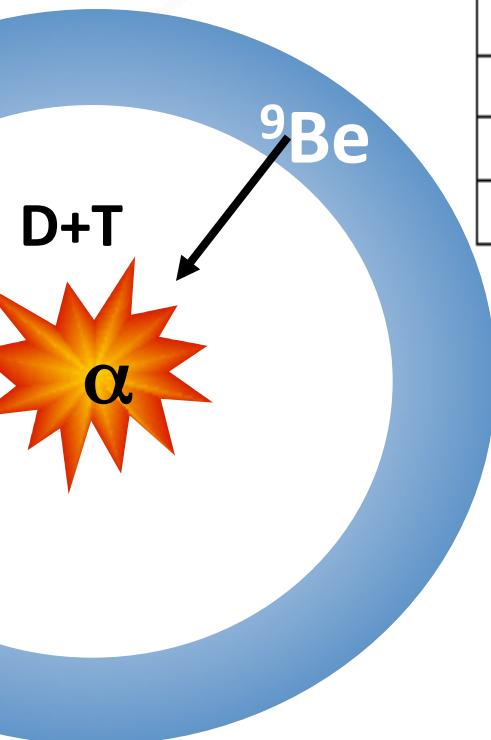
# Exploding pusher can be used to extract $^{12}\text{C}(\text{n},\gamma)$ and $\text{D}(\text{n},\gamma)$ from Ignition Capsule



# Hydro-dynamical Mixing of Ablator into Hotspot

$^9\text{Be}(\alpha, n\gamma)^{12}\text{C}$  gamma-rays  
(4.44 MeV)

HS Radius ( $\mu\text{m}$ )	0.1 $\mu\text{g}$ $^9\text{Be}$ in HS	1 $\mu\text{g}$ $^9\text{Be}$ in HS
10	$1.2 \times 10^{-4}$	$1.2 \times 10^{-3}$ $\gamma/n$
20	$3.0 \times 10^{-5}$	$3.0 \times 10^{-4}$
30	$1.35 \times 10^{-5}$	$1.35 \times 10^{-4}$



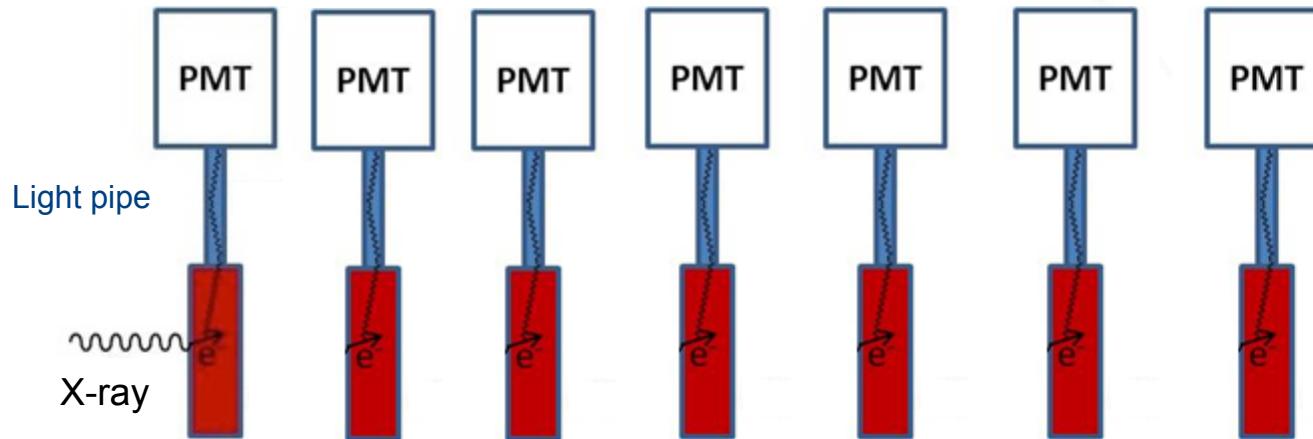
$^{10}\text{B}(\alpha, p)^{13}\text{C}$  gamma-rays  
(3.68, 3.86 MeV)

HS Radius ( $\mu\text{m}$ )	0.1 $\mu\text{g}$ $^{10}\text{B}$ in HS	1 $\mu\text{g}$ $^{10}\text{B}$ in HS
25	$1.34 \times 10^{-5}$	$1.34 \times 10^{-4}$

Compare w/  $1.2 \times 10^{-3}$   $\gamma/n$  at 100 mg/cm<sup>2</sup>  $^{12}\text{C}$

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# Aerogel Cherenkov Detector (ACD) proposed for low-energy (0.2-2.5 MeV), coarse spectroscopy



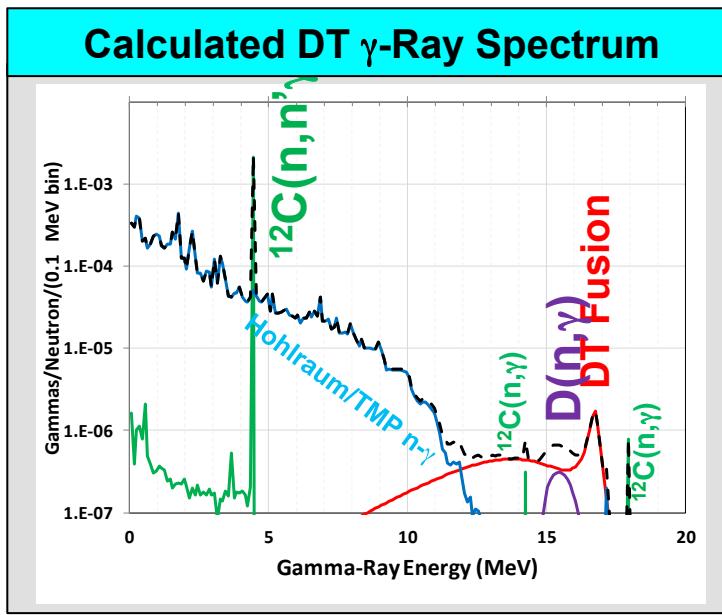
Radiator:	Aerogel-1	Aerogel-2	Aerogel-3	Aerogel-4	Aerogel-5	water	quartz
Refractive index, n:	1.015	1.03	1.05	1.07	1.12	1.33	1.5
Threshold Energy (MeV):	2.47	1.62	1.16	0.93	0.62	0.26	0.17

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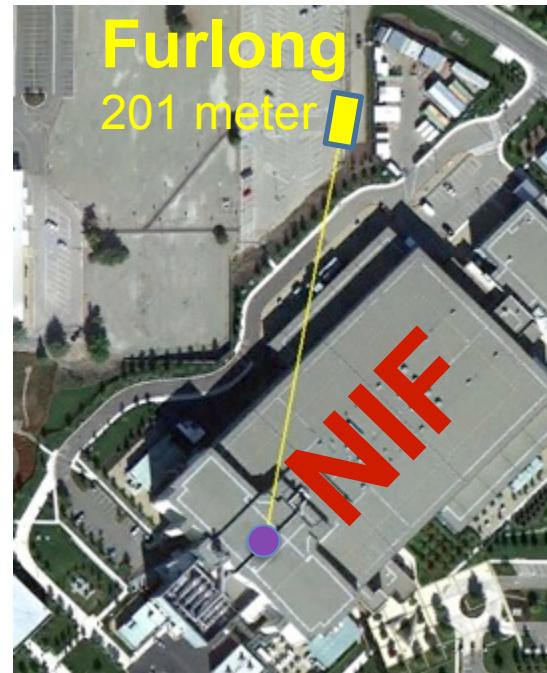
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# Options for Future Gamma Diagnostics (cont.)

- Spectroscopic detectors:
  - Curved Crystal Gamma Spec (W.Stoeffl)
    - useful for <1.5 MeV
  - “Furlong” (W.Stoeffl)
    - Single-hit, pixelated scintillator detector array
    - ~200m from TCC

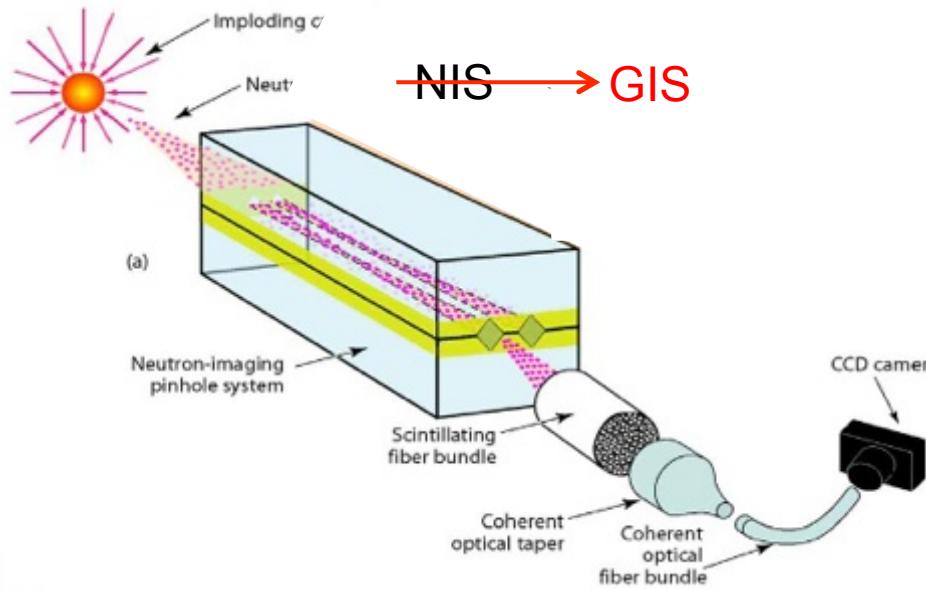


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# Gamma Imaging System (GIS)

Imaging of  $^{12}\text{C}(\text{n},\text{n}')\gamma$  would reveal ablator mass distribution at bang time

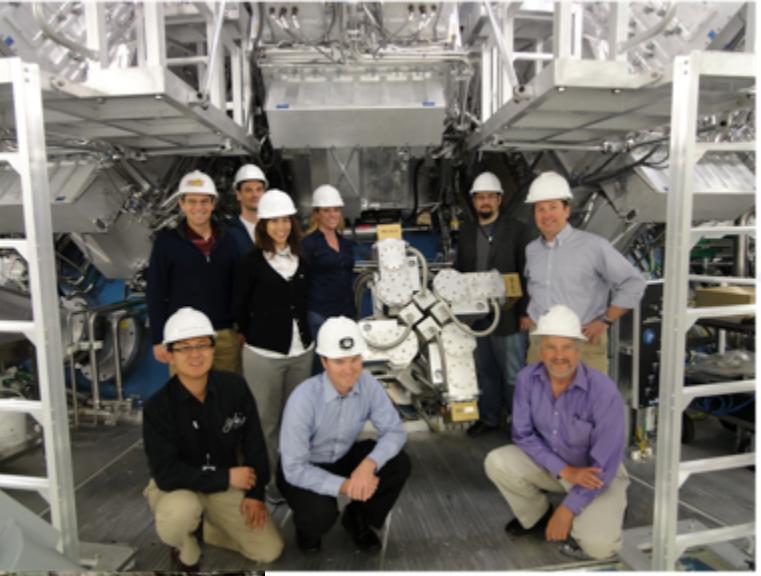
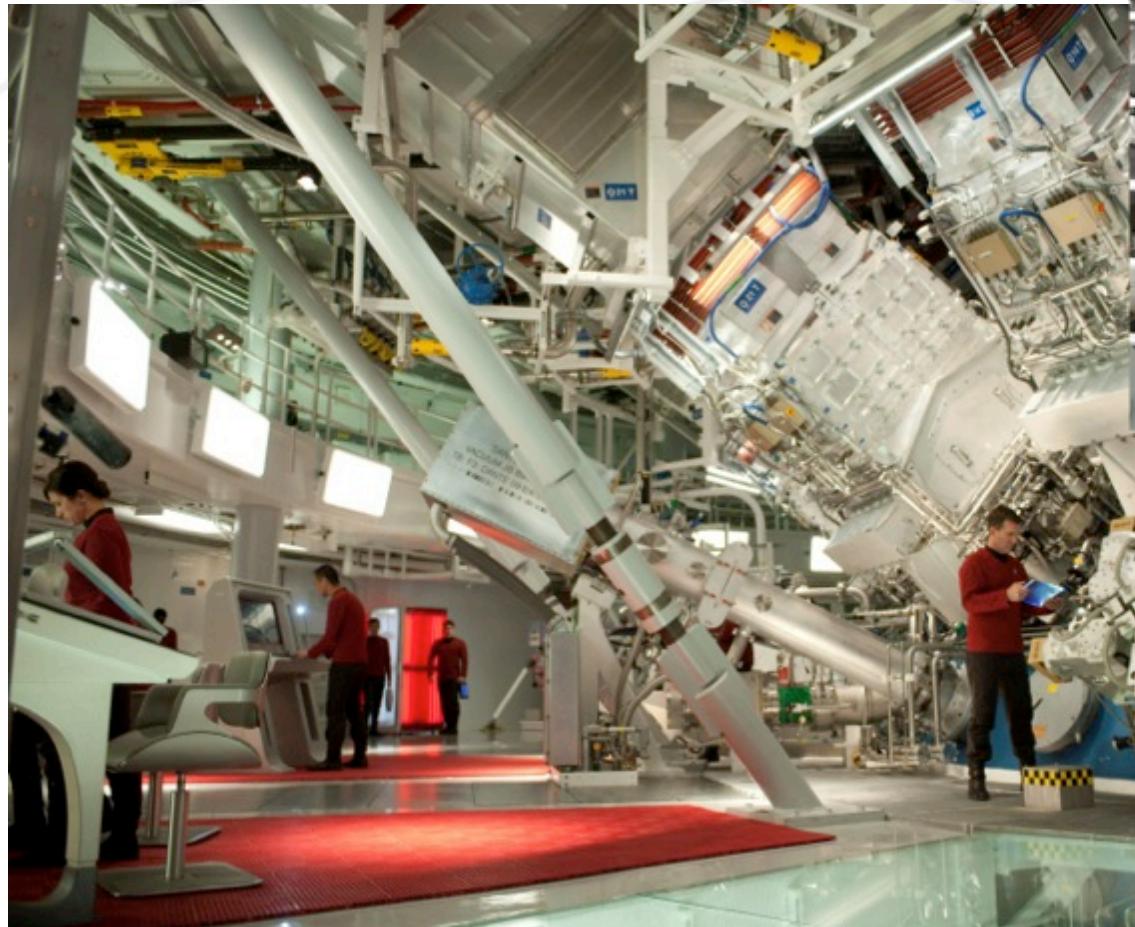


**$\gamma$ -Ray Imaging of Ablator ( $^{12}\text{C}-\gamma$ )**  
(G. Grim, D. Lemieux (U.Ariz))

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Slide 30

# GRH boldly goes...



*courtesy Scott Chambliss,  
Paramount Pictures and  
Bad Robot Productions*





## ■ Backups

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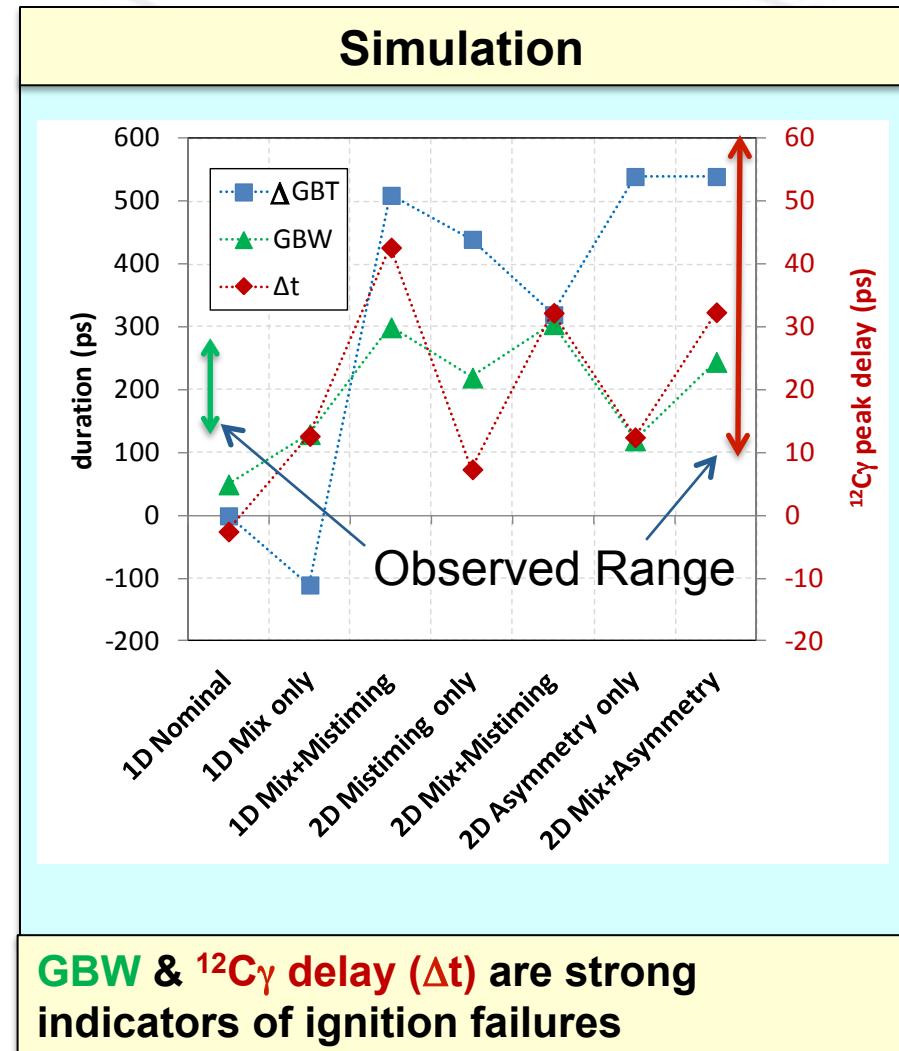


Operated by Los Alamos National Security, LLC for the U.S. Department of Energy's NNSA



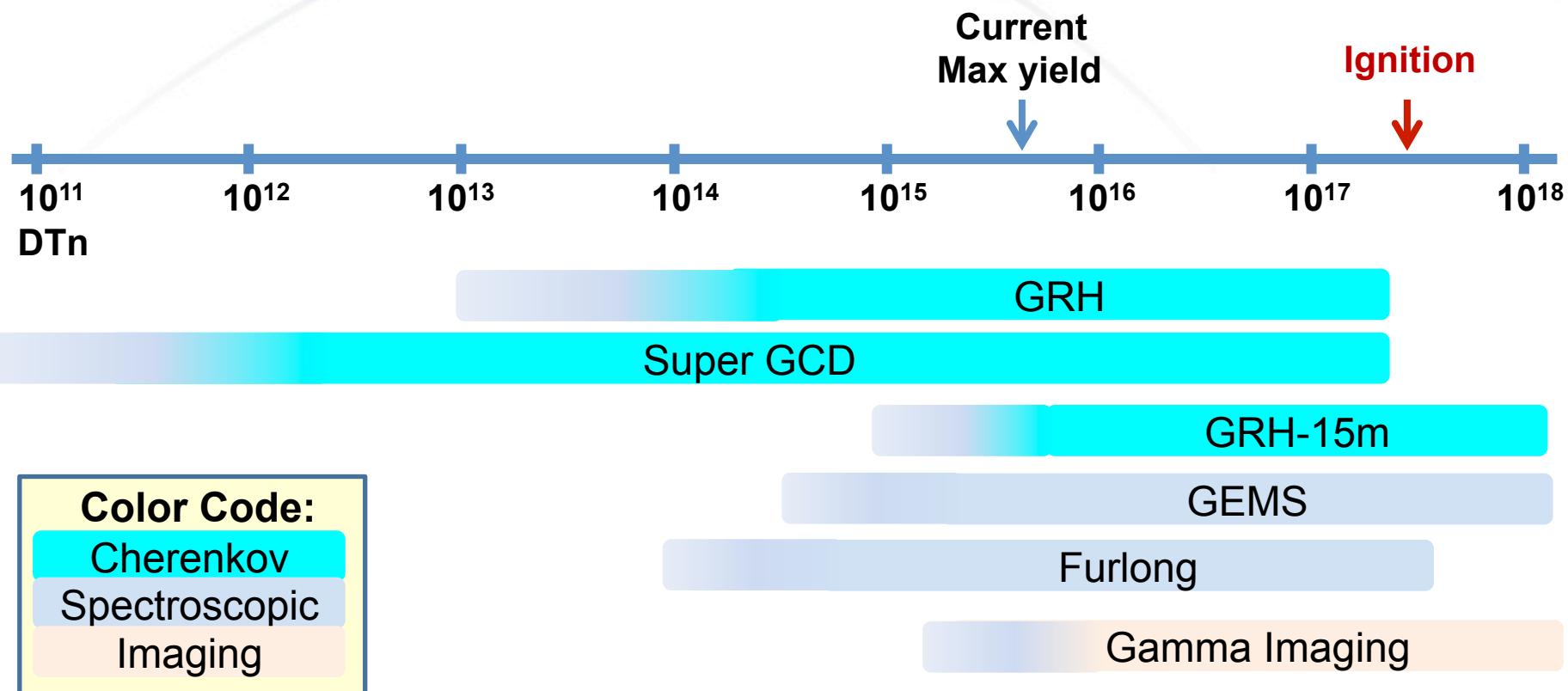
# GRH continues to inform quest for Ignition at NIF

- ~500ps late BT indicative of reduced coupling → Drive Multipliers (~85%)
  - Scatter of nBT relative to GBT indicative of core velocity
- Wide GBW indicative of failure modes during NIC
- Large & late  $^{12}\text{C}\gamma$  peak indicative of improper stagnation
- Future:  $^{12}\text{C}$  layers in  $^{13}\text{C}$  capsules for Dark Mix studies



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# NIF Yield Ranges for Gamma Diagnostics



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